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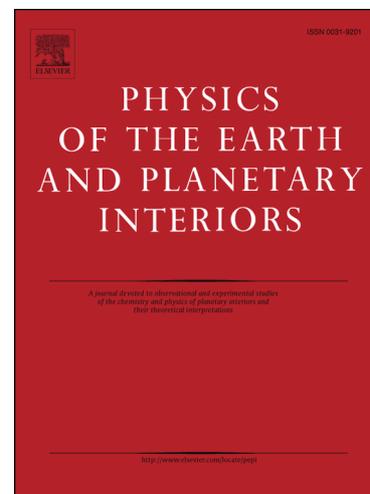
Towards constitutive equations for the deep Earth

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# Towards constitutive equations for the deep Earth

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## Abstract

A new formulation of constitutive equations for states of high compression is introduced for isotropic media, exploiting a separation between hydrostatic and deviatoric components in strain energy. The strain energy is represented as functions of strain invariants, with one purely volumetric component and the other which vanishes for purely hydrostatic deformation. This approach preserves the form of familiar equations of state through the volumetric component, but allows the addition of volume and pressure dependence of the shear modulus from the deviatoric term. A suitable shear modulus representation to accompany a Keane equation of state is demonstrated.

*Keywords:* Constitutive equations, Equations of State, Bulk Modulus, Shear Modulus, Deep Earth

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## 1. Introduction

The pressures and temperatures in the Earth's lower mantle are already high enough that properties of materials differ substantially from the ambient state. Experimental and *ab initio* computational methods have steadily improved, so that there is now substantial information available on the behaviour of the bulk modulus (K) at large compression. Recently the shear modulus (G) has also been probed for many materials of importance in the deep Earth.

The dominant representation of material behaviour for high-pressure

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25 studies is the use of the Birch-Murnaghan formulation coupled with a  
26 Mie-Grüneisen-Debye treatment of thermal effects. A systematic anisotropic  
27 formulation was provided by Stixrude and Lithgow-Bertollini (2005) from which  
28 bulk and shear moduli can be readily extracted.

29 However, many experimental studies at high compression have favoured  
30 rather different representations of bulk modulus behaviour. Thus Sakai et al.  
31 (2016) in a study of the post-perovskite phase have preferred the Keane equation  
32 of state (EOS), having tested a range of parameterisations. Yet, except for  
33 Birch-Murnaghan, there is no corresponding development for the shear modulus.

34 In this study we demonstrate that it is possible to develop an isotropic  
35 formulation of the constitutive equation between stress and strain that allows the  
36 retention of familiar equations of state for the bulk modulus, whilst including  
37 shear effects via a deviatoric component. This representation enlarges the  
38 repertoire of available ways of describing material behaviour under high pressure  
39 and temperature.

## 40 **2. Constitutive Equations**

41 A constitutive equation provides a specification of the relation between the  
42 stress tensor  $\sigma$  and a representation of strain  $\mathbf{E}$ . We will initially consider states  
43 solely under compression, and briefly introduce thermal effects in Section 3. We  
44 will follow the continuum mechanics approach and notation of Kennett and Bunge  
45 (2008), making a development in terms of strain energy  $W$ .

46 We consider a deformation from a reference state (unstressed) described by  
47 coordinates  $\xi$  to a current state described by coordinates  $\chi$ . The relation between  
48 the states is provided by the deformation gradient tensor  $\mathbf{F} = \partial\chi/\partial\xi$ , and  $J =$   
49  $\det \mathbf{F} = V/V_0$  is then the ratio of a volume element in the current state ( $V$ ) to that in  
50 the reference state ( $V_0$ ). We also introduce the displacement gradient tensor  $\mathbf{A} =$

51  $\mathbf{F} - \mathbf{I}$ , which provides a measure of the distortion introduced by the deformation.

52 In terms of  $\mathbf{F}$  and the Green strain  $\mathbf{E} = \frac{1}{2}(\mathbf{F}^T\mathbf{F} - \mathbf{I})$ , the components of the  
53 stress tensor  $\boldsymbol{\sigma}$  are given by

$$54 \quad J\sigma_{ij} = F_{ik} \frac{\partial W}{\partial F_{jk}} = F_{ik} F_{jl} \frac{\partial W}{\partial E_{kl}}, \quad (1)$$

55 where we use the Einstein summation convention of summation over repeated  
56 suffices.

57 The nature of the strain energy  $W$  thus determines the relationship between  
58 stress and strain. For an elastic material,  $W$  can be equated to the specific  
59 Helmholtz free energy  $\mathcal{F}/\rho$ , where  $\rho$  is density. In terms of specific quantities  
60 the thermodynamic relations are

$$61 \quad \rho dW = -\rho S dT + \rho_0 \sigma_{ij} dA_{ij}, \quad (2)$$

62 in terms of the displacement gradient  $\mathbf{A}$ , specific entropy  $S$  and temperature  $T$ .

63 The stress tensor  $\sigma_{ij}$  can be derived from  $\mathcal{F}$  as

$$64 \quad \sigma_{ij} = \frac{\rho}{\rho_0} \frac{\partial \mathcal{F}}{\partial A_{ij}} \quad (3)$$

65 since the volume ratio  $J$  can also be written as  $J = \rho_0/\rho$ .

66 The most complete current formulation of such a constitutive equation is that  
67 by Stixrude and Lithgow-Bertelloni (2005), based on the earlier work of Birch and  
68 Murnaghan. This employs a Taylor series expansion of the Helmholtz free energy  
69 about the reference state in terms of the Eulerian strain tensor  $\mathbf{e} = \frac{1}{2}(\mathbf{I} - [\mathbf{F}\mathbf{F}^T]^{-1})$ .

70 The volume transformation

$$71 \quad \left(\frac{\rho}{\rho_0}\right)^2 = J^{-2} = \det[2\mathbf{e} - \mathbf{I}]. \quad (4)$$

72 The Helmholtz Free energy is then written as a power series in the Eulerian strain

$$73 \quad \mathcal{F} = V_0 \sum_i B_i \mathbf{e}^i, \quad (5)$$

74 this Birch-Murnaghan formulation is commonly taken to 3 or 4 terms.

75 By examining local perturbations from a stressed state the elastic moduli  $K$ ,  $G$   
76 can be extracted. The choice of Eulerian strain markedly reduces the influence  
77 of the third-order term in strain in (5). The third-order representation does not  
78 involve any second derivatives of moduli. When coupled with a representation  
79 of thermal pressures with a Debye-Mie-Grüneisen form this provides a complete  
80 system for characterising states with moderate pressure (as in Section 3).

81 The disadvantage of this approach is that it essentially extrapolates from  
82 low pressure to higher pressures, depending strongly on the gradients of the  
83 moduli ( $K'_0$ ,  $G'_0$ ) in the reference state. The situation is improved if high-pressure  
84 information is available for a material, but even then differences can arise from  
85 the way in which the inversion for the set of mechanical and thermal parameters  
86 is conducted. Kennett and Jackson (2009) have demonstrated that a full nonlinear  
87 inversion can be effective, and provide both uncertainty estimates and information  
88 about cross-coupling between parameters.

### 89 2.1. Equations of State

90 In many situations a reduced form of the constitutive equation is employed  
91 relating volume  $V$ , pressure  $p$  and temperature  $T$ . Such *equations of state* (EOS)  
92 can only describe the behaviour of the bulk modulus ( $K$ ). A number of different  
93 formulations have been used to fit experimental data on material properties at high  
94 pressure, and can be written in terms of the density ratio  $\chi = \rho/\rho_0 = V_0/V = J^{-1}$ .

95 The ‘cold’ part of equations of state provides a specification of the pressure  
96  $p$  as a function of volume  $p(V)$  or, equivalently, density ratio  $p(\chi)$ . The bulk  
97 modulus  $K$  can be extracted from the expressions for the pressure in the EOS from  
98  $K = -V(\partial p/\partial V)_T = \chi(\partial p/\partial \chi)$ . A further differentiation extracts the pressure  
99 derivative  $K' = (\partial K/\partial p)_T = \chi(\partial K/\partial p)/K$ .

100 The Vinet-Rydbberg-Morse EOS (Vinet et al., 1987) is based on an atomic force  
101 model, with pressure represented as

$$102 \quad p = 3K_0x^{2/3}[1 - x^{-1/3}] \exp\{\frac{3}{2}(K'_0 - 1)[1 - x^{-1/3}]\}, \quad (6)$$

103 where  $K_0$  is the bulk-modulus at ambient conditions, and  $K'_0 = [\partial K/\partial p]_0$  is its  
104 pressure derivative. The bulk modulus as a function of the density ratio  $x$  is then

$$105 \quad K = K_0x^{2/3} [2 + (\zeta - 1)x^{-1/3} - \zeta x^{-2/3}] \exp\{\zeta[1 - x^{-1/3}]\}, \quad (7)$$

106 where  $\zeta = \frac{3}{2}(K'_0 - 1)$ .

107 Poirier and Tarantola (1998) used a similar development to the  
108 Birch-Murnaghan approach, but employed logarithmic strain, which gives a  
109 more rapid convergence. To second order, the pressure is

$$110 \quad p = K_0x [\ln x + \frac{1}{2}(K'_0 - 2)(\ln x)^2]. \quad (8)$$

111 Although originally derived using logarithmic strain, the Poirier-Tarantola EOS  
112 (8) can be recognised as simply a function of the strain invariant  $x = 1/J$ . The  
113 associated representation of the bulk modulus is

$$114 \quad K = K_0x [1 + (K'_0 - 1) \ln x + \frac{1}{2}(K'_0 - 2)(\ln x)^2]. \quad (9)$$

115 Stacey and Davis (2004) advocate the use of the Keane (1954) EOS for deep  
116 Earth studies because it links to properties at (nominal) infinite pressure:

$$117 \quad p = K_0 \left[ \frac{K'_0}{K'^2_\infty} [x^{K'_\infty} - 1] - \left( \frac{K'_0}{K'_\infty} - 1 \right) \ln x \right]. \quad (10)$$

118 Thermodynamic arguments suggest a lower bound on  $K'_\infty$  of 5/3. The Keane EOS  
119 can be regarded as an interpolant rather than just an extrapolant, though the high  
120 pressure limit enters as a parameter in fitting. The Keane representation of the  
121 bulk modulus has a rather simple form,

$$122 \quad K = K_0 \left[ 1 + \frac{K'_0}{K'_\infty} (x^{K'_\infty} - 1) \right]. \quad (11)$$

123 Each EOS should be regarded as a parametric representation of behaviour,  
 124 and thus when different expressions are used to fit the same experimental data the  
 125 values obtained for  $K_0$ ,  $K'_0$  will be similar but not identical (see, e.g., Sakai et al.  
 126 2016).

127 None of these equations of state have any associated shear moduli. Further,  
 128 unlike the Birch-Murnaghan expansion, none has any obvious extensions to tensor  
 129 form that would allow extraction of shear properties.

### 130 2.2. Isotropic Constitutive Equations

131 If we concentrate attention on just the bulk modulus ( $K$ ) and shear modulus  
 132 ( $G$ ) we can describe behaviour in terms of isotropic constitutive equations. The  
 133 important materials in the deep Earth, e.g. bridgmanite and ferro-periclase, are  
 134 intrinsically anisotropic at the crystal level. Nevertheless, the properties of  
 135 aggregates can be adequately described in isotropic terms, as is commonly used.

136 For an isotropic medium, the strain energy  $W$  can be represented as a function  
 137 of invariants of the strain measures (Spencer, 1980). An extensive development  
 138 has been made for large deformation in rubber-like materials in tension, whereas  
 139 we need results for strong compression.

140 The deformation gradient  $\mathbf{F}$  can be written in terms of a stretching component  
 141 and a rotation in two ways

$$142 \quad \mathbf{F} = \mathbf{R}\mathbf{U} = \mathbf{V}\mathbf{R} \quad (12)$$

143 where  $\mathbf{U}^2 = \mathbf{F}^T\mathbf{F}$  and  $\mathbf{V}^2 = \mathbf{F}\mathbf{F}^T$ .  $\mathbf{U}$ ,  $\mathbf{V}$  have the same eigenvalues, the principal  
 144 stretches  $\lambda_1, \lambda_2, \lambda_3$ , but the principal axes vary in orientation by the rotation  $\mathbf{R}$ .

145 The useful invariants of  $\mathbf{U}$ ,  $\mathbf{V}$  are

$$146 \quad J^2 = \lambda_1^2 \lambda_2^2 \lambda_3^2 = \det \mathbf{U}^2, \quad (13)$$

147 a purely hydrostatic term, representing changes in volume, and

$$148 \quad L = J^{-2/3}[\lambda_1^2 + \lambda_2^2 + \lambda_3^2] = J^{-2/3} \text{tr} \mathbf{A}^2 = J^{-2/3} \text{tr} \mathbf{U}^2. \quad (14)$$

149 which concentrates on the deviatoric aspects of deformation.

150 In such an isotropic medium the principal axes of the stress tensor  $\boldsymbol{\sigma}$  align with  
 151 those of  $\mathbf{V}$  (the Eulerian triad), whereas the principal axes of  $\mathbf{U}$  and  $\mathbf{E}$  are rotated  
 152 by  $\mathbf{R}$  (the Lagrangian triad). In terms of the principal stretches we can recast (1)  
 153 in the form of an expression for the  $r$ th principal stress

$$154 \quad \sigma_r = \frac{1}{J} \lambda_r \frac{\partial W}{\partial \lambda_r}, \quad \text{no sum on } r, \quad (15)$$

155 whilst recognising the rotation between the principal directions of the elements on  
 156 the left- and right-hand sides of the equation (15).

157 Now consider a strain energy function  $W$  as a function of the stretch invariants  
 158  $J$ ,  $L$  with two independent volume terms  $\Phi(J)$  and  $\Psi(J)$ :

$$159 \quad W = \Phi(J) + \{L - 3\} \Psi(J); \quad (16)$$

160 incorporating a direct volume dependence in  $\Phi(J)$  and a deviatoric component in  
 161 the second term. For *purely hydrostatic compression*  $\lambda_1 = \lambda_2 = \lambda_3 = \bar{\lambda}$ ,  $J = \bar{\lambda}^3$   
 162 and  $\{L - 3\} = \bar{\lambda}^{-2} 3 \bar{\lambda}^2 - 3 = 0$ , so that the deviatoric term  $\{L - 3\} \Psi(J) = 0$ .

163 As detailed in Appendix A1, the  $r$ th principal stress derived from the strain  
 164 energy form (16) is

$$165 \quad \sigma_r = \frac{\partial \Phi}{\partial J} + \frac{2}{J^{5/3}} \left[ \lambda_r^2 - \frac{1}{3} \text{tr} \boldsymbol{\Lambda}^2 \right] \Psi(J) + \{L - 3\} \frac{\partial \Psi}{\partial J}. \quad (17)$$

166 For a hydrostatic state, when  $L - 3 = 0$ , the dependence on  $\Psi(J)$  vanishes and so

$$167 \quad \sigma_r = -p = \frac{\partial \Phi}{\partial J}, \quad (18)$$

168 i.e., isotropic stress independent of the form of  $\Psi(J)$ . Each of the equation of  
 169 state expressions in (6), (8) and (10) correspond to a specification of  $\partial \Phi / \partial J$ , even  
 170 though the original derivations did not explicitly use the strain invariant.

171 The full stress tensor

$$172 \quad \boldsymbol{\sigma} = \mathbf{R} \left\{ \left( \frac{\partial \Phi}{\partial J} + \{L - 3\} \frac{\partial \Psi}{\partial J} \right) \mathbf{I} + \frac{2}{J^{5/3}} \left[ \mathbf{U}^2 - \frac{1}{3} \text{tr}(\mathbf{U}^2) \mathbf{I} \right] \Psi(J) \right\} \mathbf{R}^T, \quad (19)$$

173 with strong simplification in the hydrostatic case when  $\{L - 3\} = 0$  and the term  
174 in  $\Psi$  vanishes to:

$$175 \quad -p\mathbf{I} = \frac{\partial\Phi}{\partial J}\mathbf{I}. \quad (20)$$

176 The representation (16) with a separation of hydrostatic and deviatoric parts was  
177 suggested by neo-Hookean equations for rubbers, but now includes a volume  
178 (density) modulation of  $\{L - 3\}$  through  $\Psi(J)$  to allow for strong compression.

179 The elastic moduli as a function of density (and hence pressure) can be  
180 extracted from the stress tensor in the form (19) by making a first order expansion  
181 about a hydrostatic compressed state with  $\lambda_r = \bar{\lambda}(1 + e_r)$ . Then, e.g.,  $J =$   
182  $\bar{\lambda}^3(1 + \text{tr}\{\mathbf{e}\}) + O(e^2)$ .

183 The details of the first order expansion about the hydrostatic state are presented  
184 in Appendix A2. The  $r$ th principal stress in terms of  $\mathbf{e}$  reduces to:

$$185 \quad \sigma_r = -p + J \frac{\partial^2\Phi}{\partial J^2} \text{tr}\{\mathbf{e}\} + \frac{2}{J} \Psi(J) [e_r - \frac{1}{3} \text{tr}\{\mathbf{e}\}]. \quad (21)$$

186 We can recognise the elastic moduli by comparison with the standard form for  
187 isotropic media

$$188 \quad \sigma_r = -p + K \text{tr}\{\mathbf{e}\} + G [e_r - \frac{1}{3} \text{tr}\{\mathbf{e}\}] \quad (22)$$

189 so that we have:

$$\begin{aligned} \text{Bulk Modulus} \quad K &= J \partial^2 \Phi(J) / \partial J^2, \\ \text{Shear Modulus} \quad G &= 2\Psi(J) / J. \end{aligned} \quad (23)$$

190 We have thus demonstrated that it is possible to retain existing EOS  
191 representations of the bulk modulus  $K$  with a suitable specification of  $\Phi(J)$ , but  
192 to attach shear dependence through a new function of volume (density)  $\Psi(J)$ . In  
193 terms of the density ratio  $\alpha$  the shear modulus  $G$  and shear wavespeed  $\beta$  take the  
194 form:

$$195 \quad G = \frac{2}{J} \Psi(J) = 2\alpha \Psi(\alpha), \quad \beta^2 = \frac{G}{\rho} = 2\rho_0 \Psi(\alpha). \quad (24)$$

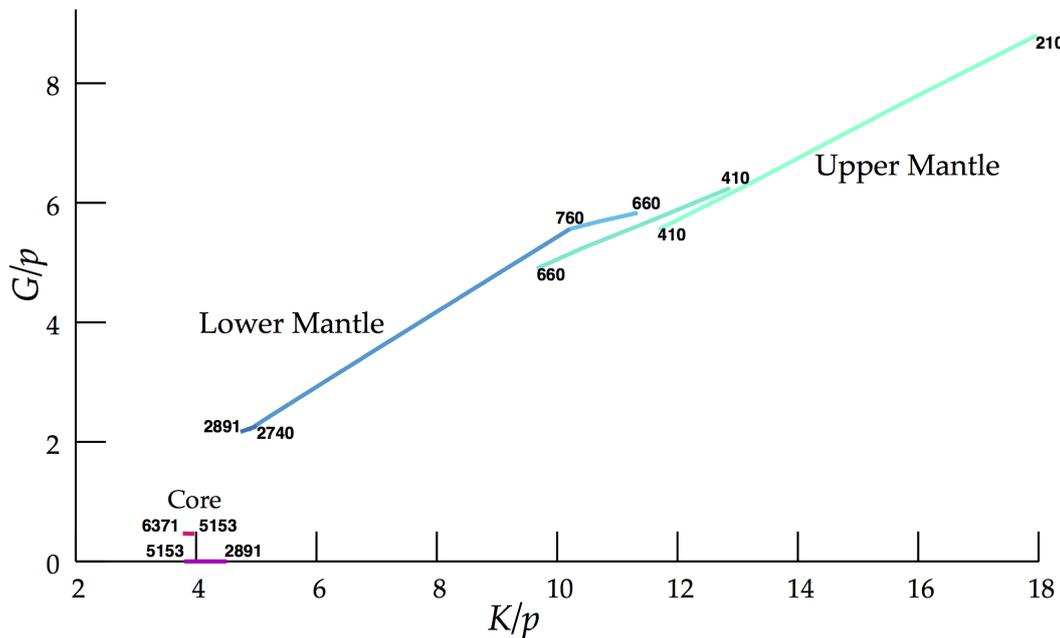


Figure 1: Multiple linear segments for the relation between  $G/p$  and  $K/p$  for different parts of the *ak135* Earth model (Kennett et al., 1995). The depths for each segment are indicated. Multiple depths indicate the presence of major discontinuities.

196 It is thus possible to capture the volume dependence of the shear modulus in a  
 197 simple form. But, since the pressure equation relates only to  $K$ , the pressure  
 198 derivative  $G'$  will be coupled to  $K, K'$ .

### 199 2.3. Building constitutive equations

200 The treatment of Section 2.2 demonstrates that we can specify pressure and  
 201 bulk modulus behaviour as a function of compression through a strain energy  
 202 contribution  $\Phi(J)$ , with the description of the shear modulus to be assigned  
 203 through a separate function  $\Psi(J)$ .

204 The group of equations of state considered in Section 2.1 already provide  
 205 different representations of pressure and bulk modulus, and can thus be used  
 206 directly. But, how then should we link in shear properties?

207 For current Earth Models, empirical relations of the form

$$208 \quad G = aK - bp, \quad (25)$$

209 provide good piecewise fits to segments of radial Earth structure (Figure 1), such  
 210 as the entire lower mantle. Different coefficients  $a$ ,  $b$  describe the behaviour  
 211 of the various segments, indicated by different tones in Figure 1. At major  
 212 discontinuities such as the ‘410 km’ and ‘660 km’ discontinuities or the inner-core  
 213 boundary, the moduli  $K$ ,  $G$  show discontinuous increases with depth at constant  
 214 pressure so that segments can overlap.

215 The empirical relation (25) suggests that we should seek functional forms for  
 216 the representation of  $G(x)$  that incorporate the dependencies on density ratio of  
 217 both bulk modulus and pressure for any particular equation of state. Thus the  
 218 functional form of  $\Psi(J)$  combines elements from  $\Phi'(J)$ ,  $\Phi''(J)$ .

219 Although the Vinet-Rydberg-Morse equation of state is frequently effective  
 220 in representing bulk modulus behaviour, it is based on a central potential model  
 221 that does not readily relate to shear. We therefore demonstrate how a shear  
 222 counterpart to the Keane EOS can be constructed. We combine the suite of  
 223 functional dependencies from (10) and (11) to suggest a representation

$$224 \quad G(x) = G_0 \left( A \ln x + Bx^{K'_\infty} + (1 - B) \right), \quad (26)$$

225 with pressure derivative

$$226 \quad G'(x) = \frac{\partial G}{\partial p} = \frac{x}{K} \frac{\partial G}{\partial x} = \frac{G_0}{K(x)} \left( A + BK'_\infty x^{K'_\infty} \right), \quad (27)$$

227 where  $K(x)$  is given by (11). The constant  $A$  is unconstrained by the initial  
 228 condition on the modulus, but can be extracted from  $G'_0$  as

$$229 \quad A = G'_0 \left( \frac{K_0}{G_0} \right) - BK'_\infty. \quad (28)$$

230 The expression for the shear modulus is thus strongly linked to that of the bulk  
 231 modulus, but has three independent parameters  $G_0$ ,  $G'_0$  and  $B$ . In a similar way

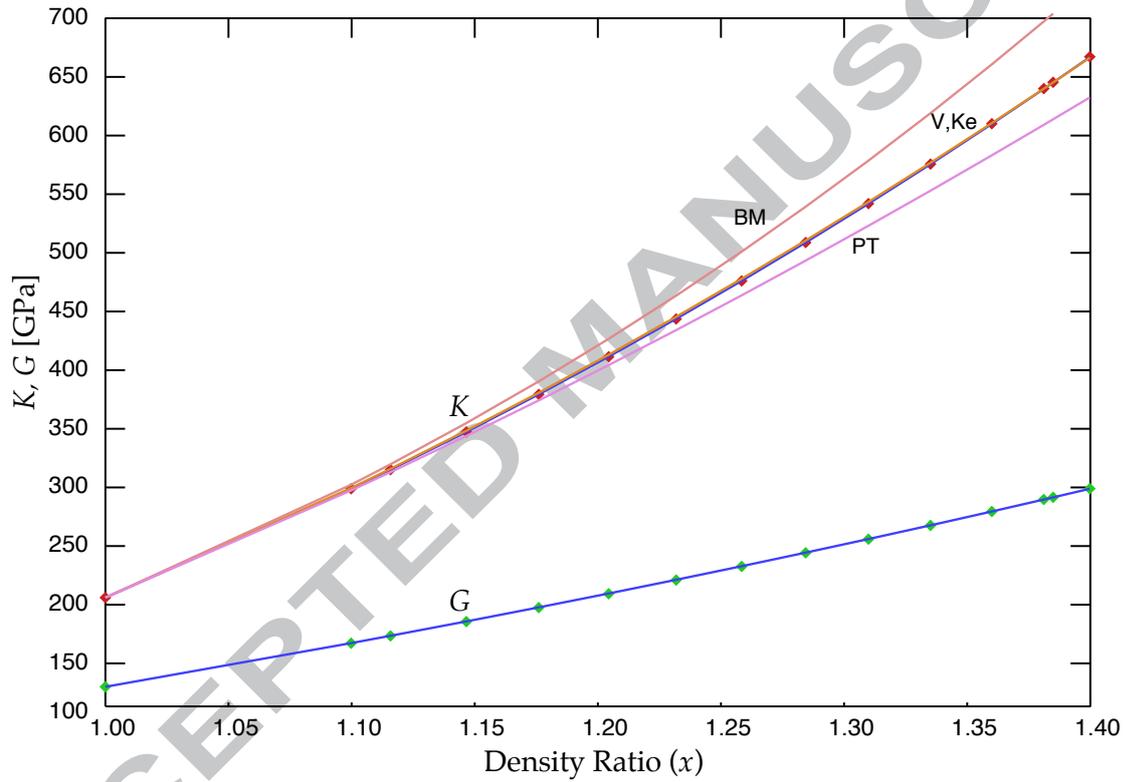


Figure 2: Bulk and shear modulus for the lower mantle as a function of the density ratio  $x$ , with EOS fits using the same values of  $K_0$ ,  $K'_0$  and a fit to the shear modulus using (26). BM: Birch-Murnaghan; PT: Poirier-Tarantola; V: Vinet-Rydberg-Morse, Ke, Keane.

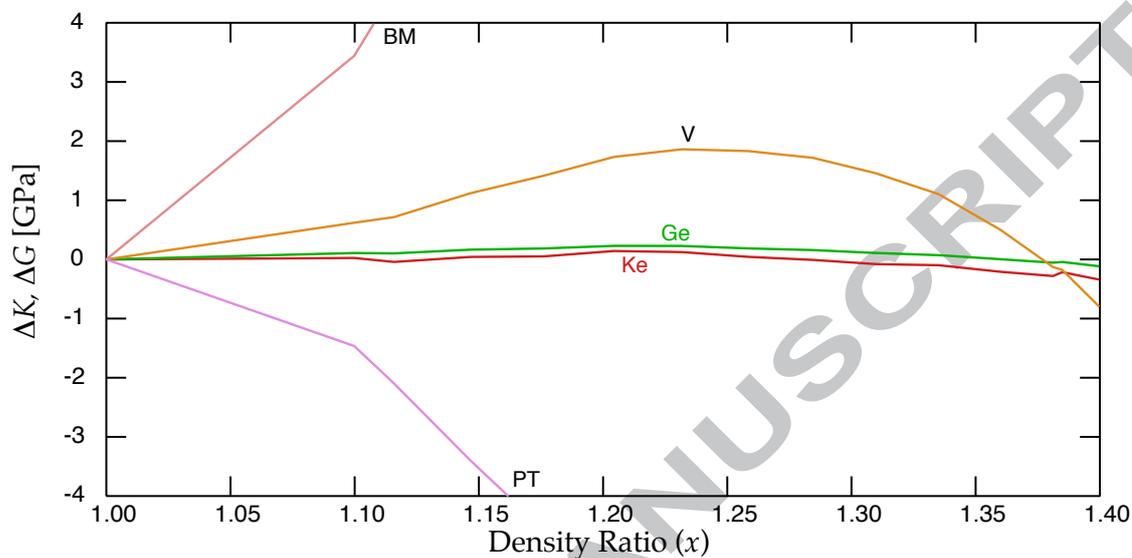


Figure 3: Deviations of fits to Bulk and shear modulus for the lower mantle as a function of the density ratio  $x$ . BM: Birch-Murnaghan; PT: Poirier-Tarantola; V: Vinet-Rydberg-Morse, Ke, Keane; Ge: fit to the shear modulus using (26).

232 to the three parameter fit for the bulk modulus ( $K_0, K'_0, K'_\infty$ ) this triad of shear  
 233 parameters provides considerable flexibility in fitting data.

234 The differences between the various styles of representation of dependence  
 235 on compression  $x$  only become evident for conditions corresponding to the lower  
 236 mantle and deeper. We use the  $K, G$  values from Table 1 of Stacey and Davis  
 237 (2004), ignoring temperature effects, as a sample with a wide span of density  
 238 ratios (Figure 2). We compare the suite of equations of state with the same  
 239 nominal  $K_0$  value (206.06) and  $K'_0$  (4.2), and show how we can use the shear  
 240 equation (26), linked to the Keane EOS, to fit the  $G$  distribution.

241 For density ratios up to 1.10 there is essentially no difference in the values  
 242 from the third-order Birch-Murnaghan form or any of the other EOS. As  
 243 compression increase the results diverge. The Vinet-Rydberg-Morse and Keane  
 244 results fit the data points well, but the Birch-Murnaghan and Poirier-Trantola

245 forms deviate significantly for larger  $x$ . In each case, adjustment of the values  
246 of  $K'_0$  can improve the fit, though not over the full range of compression.

247 The Vinet-Rydberg-Morse results provide a good two-parameter fit to the  
248 specified  $K(x)$  values with deviations less than 2 GPa ( $\sim 0.4\%$ ), as can be  
249 seen in Figure 3 that compares the deviations from the specified values. The  
250 three-parameter fit with the Keane EOS ( $K'_\infty = 2.575$ ) is even better ( $< 0.05\%$ ).  
251 With this same set of specified  $K_0$ ,  $K'_0$ , and  $K'_\infty$ , the shear representation (26) is  
252 readily tuned to match the  $G(x)$  exceptionally well ( $G_0=130.02$ ,  $G'_0=1.745$ ,  $B =$   
253  $0.72$ ), as can be seen in Figure 3.

254 This example demonstrates that linked bulk and shear modulus representations  
255 can be satisfactorily developed exploiting the functional dependencies suggested  
256 by the empirical relation (25). The need is strongest for high compression, and  
257 shear information is beginning to become available in this regime as experimental  
258 techniques improve.

### 259 3. Mie-Grüneisen-Debye thermal contribution

260 In order to construct a full constitutive equation we need to include thermal  
261 effects as well as those associated with deformation. This can readily be done by  
262 including an additional contribution to the specific Helmholtz Free Energy:

$$263 \mathcal{F}(V, T) = \mathcal{F}_C(V, 0) + \mathcal{F}_D(V, T), \quad (29)$$

264 combining a ‘cold’ part  $\mathcal{F}_C(V, 0)$  and a ‘warm’ part  $\mathcal{F}_D(V, T)$  as in Stixrude and  
265 Lithgow-Bertelloni (2005). Then the contributions to the *elastic moduli* can be  
266 thought of in terms of trajectories in an  $M$ ,  $T$ ,  $p$  space (Figure 4).

267 The contribution from lattice vibrations can be well represented by the Debye  
268 form

$$269 E_D(T) = 9nRT \left( \frac{T}{\theta} \right)^3 \int_0^{\theta/T} d\xi \frac{\xi^3}{\exp(\xi) - 1}, \quad (30)$$

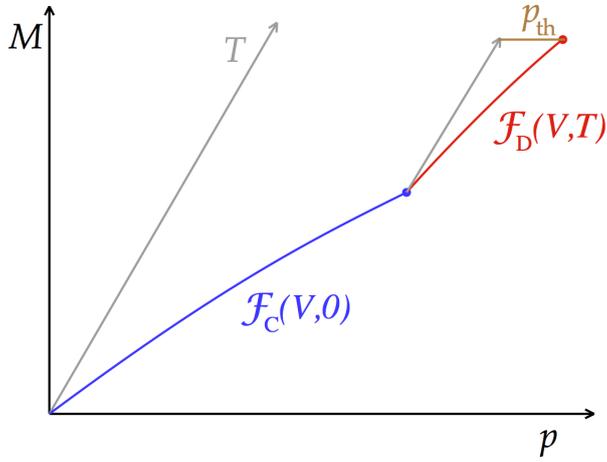


Figure 4: Inclusion of thermal stress as an additional component in  $M$ ,  $T$ ,  $p$  space.

270 where  $n$  is the number of atoms in the unit cell,  $R$  is the gas constant. This  
 271 simple form is effective since the net thermal effect is not sensitive to the details  
 272 of the electron distribution (Stixrude and Lithgow-Bertelloni, 2005). The lattice  
 273 vibrations add an additional thermal component to the pressure

$$274 \quad p(V, T) = p_C(V, 0) + \frac{\gamma_D}{V} E_D(V, T) \quad (31)$$

275 where  $\gamma_D$  is the Grüneisen parameter.

276 The temperature dependence of pressure is given by

$$277 \quad \left[ \frac{\partial p}{\partial T} \right]_V = \alpha K_T(V, T) = \gamma_D \frac{C_V}{V} \quad (32)$$

278 in terms of the isothermal bulk modulus  $K_T = -V[\partial p / \partial V]_T$  and the thermal  
 279 expansion coefficient  $\alpha = (1/V)(\partial V / \partial T)_p$ . The other thermal parameters are

280 based on the quasi-harmonic approximation

$$281 \quad \gamma = -\frac{d \ln v}{d \ln V} = \gamma_D, \quad q = \frac{d \ln \gamma}{d \ln V}. \quad (33)$$

282 with also  $\eta_s$  as the shear strain derivative of the Grüneisen parameter  $\gamma$ . The  
 283 adiabatic bulk modulus  $K_S$  is then given by

$$284 \quad K_S = K_T(1 + \alpha \gamma T), \quad \left[ \frac{\partial p}{\partial T} \right]_S = \frac{K_S}{\gamma T}, \quad (34)$$

285 Stixrude and Lithgow-Bertelloni (2005) provide convenient forms for  
286 the strain dependence of the Grüneisen parameter  $\gamma$  and  $\eta_s$ . Alternative  
287 representations such as that due to Al'tshuler et al (1987) can also be employed.

#### 288 **4. Discussion and Conclusions**

289 By casting the strain energy function  $W$  for an isotropic medium as a function  
290 of strain invariants in a form that allows complete separation between hydrostatic  
291 and deviatoric components, we have been able to retain familiar forms for  
292 equations of state with the addition of a full description of shear. The functional  
293 form of the shear modulus as a function of volume does not depend on the bulk  
294 modulus, but the representations are coupled through pressure dependence and  
295 pressure derivatives.

296 The current approach thus provides a functional alternative to the use of the  
297 Birch-Murnaghan finite-strain formulation for shear, with considerable flexibility  
298 available in the description of shear behaviour. Further we do not need to impose  
299 adiabatic corrections to the shear modulus. The linear dependence between  $K/p$   
300 and  $G/p$  for current Earth models, suggests that the elements included in the  
301 volume dependence of the shear modulus should be similar to those used for the  
302 bulk modulus and pressure. We have shown that a shear counterpart to the Keane  
303 EOS can be constructed exploiting these dependencies, exploiting the constraints  
304 from bulk-modulus fitting. There are no shear analogues of the thermodynamic  
305 constraints on the properties of the bulk modulus at extreme compression.

306 For many materials the range of conditions accessible to experiment is  
307 still limited, and so properties at high compression will commonly require  
308 extrapolation. It is just in this high compression regime that, as noted by Poirier  
309 and Tarantola (1998), the differences in constitutive relations become important  
310 (Figures 2, 3). By bringing in constraints from very high pressures the problem

311 is converted to a more suitable interpolation, even though this also involves  
 312 parameter fitting. With the addition of linked shear representations we can expect  
 313 to improve the description of very high pressure phases, and hopefully understand  
 314 the complex variations of shear wavespeed in seismic images of the lowermost  
 315 mantle.

316 The approach we have employed to link in a shear component to the  
 317 constitutive equation is specific to the isotropic situation, and there is no  
 318 immediate generalisation to the fully anisotropic case. Yet, the functional form  
 319 of the constitutive equation (19) suggests that there may be merit in seeking  
 320 anisotropic tensor forms in which stress depends on multiple measures of strain  
 321 such as those proposed by Hill (1968). The family of Seth-Hill tensors have  
 322 the same first order expansion, but different dependence on finite strain that may  
 323 be exploited to produce suitable general constitutive relations.

## 324 **Appendix A. Appendix: mathematical derivations**

### 325 *Appendix A.1. Principal stress relations*

326 For the strain energy

$$327 \quad W = \Phi(J) + \{L - 3\}\Psi(J), \quad \text{with } \{L - 3\} = \left\{ \frac{1}{J^{-2/3}}(\lambda_1^2 + \lambda_2^2 + \lambda_3^2) - 3 \right\}, \quad (\text{A.1})$$

328 the  $\sigma_r$  principal stress takes the form

$$329 \quad \sigma_r = \frac{1}{J} \lambda_r \frac{\partial W}{\partial \lambda_r} = \frac{\lambda_r}{J} \frac{\partial J}{\partial \lambda_r} \frac{\partial \Phi}{\partial J} + \frac{\lambda_r}{J} \left[ \frac{\partial}{\partial \lambda_r} \{L - 3\} \Psi(J) + \{L - 3\} \frac{\partial J}{\partial \lambda_r} \frac{\partial \Psi}{\partial J} \right]. \quad (\text{A.2})$$

330 Now

$$\frac{\lambda_r}{J} \frac{\partial J}{\partial \lambda_r} = 1, \quad (\text{A.3})$$

$$\frac{\lambda_r}{J} \frac{\partial}{\partial \lambda_r} \{L - 3\} = \frac{1}{2J^{-2/3}} \left( 2\lambda_r^2 - \frac{2}{3}(\lambda_1^2 + \lambda_2^2 + \lambda_3^2) \right). \quad (\text{A.4})$$

331 Thus the  $\sigma_r$  principal stress takes the form

$$332 \quad \sigma_r = \frac{\partial\Phi}{\partial J} + \frac{2}{J^{-5/3}} [\lambda_r^2 - \frac{1}{3}\text{tr}\mathbf{\Lambda}^2] + \{L - 3\} \frac{\partial\Psi}{\partial J}. \quad (\text{A.5})$$

333 For an isotropic medium the principal stress align with the Eulerian triad, and the  
334 principal stretches with the Eulerian triad so that the full stress tensor takes the  
335 form

$$336 \quad \boldsymbol{\sigma} = \mathbf{R} \left\{ \left( \frac{\partial\Phi}{\partial J} + \{L - 3\} \frac{\partial\Psi}{\partial J} \right) \mathbf{I} + \frac{2}{J^{5/3}} [\mathbf{U}^2 - \frac{1}{3}\text{tr}(\mathbf{U}^2)\mathbf{I}] \Psi(J) \right\} \mathbf{R}^T, \quad (\text{A.6})$$

337 with rotation by  $\mathbf{R}$ .

338 For a hydrostatic deformation the stretches are equal,  $\lambda_1 = \lambda_2 = \lambda_3 = \bar{\lambda}$  and  
339 so  $J = \bar{\lambda}^3$ ,  $\lambda_r^2 - \frac{1}{3}\text{tr}\mathbf{\Lambda}^2 = 0$  and  $L - 3 = 0$ . The isotropic stress then reduces to

$$340 \quad -p\mathbf{I} = \frac{\partial\Phi}{\partial J} \mathbf{I} \quad (\text{A.7})$$

341 in terms of pressure  $p$ .

#### 342 *Appendix A.2. Derivation of moduli*

343 Consider making a first order perturbation about a hydrostatic compressed  
344 state with  $\lambda_r = \bar{\lambda}(1 + e_r)$ , so that  $J = \bar{\lambda}^3(1 + \text{tr}\{\mathbf{e}\}) + O(e^2)$ . Then the  $\sigma_r$  principal  
345 stress from (A.4) takes the form

$$\begin{aligned} 346 \quad \sigma_r = & \frac{\partial\Phi}{\partial J} + \text{tr}\{\mathbf{e}\}J \frac{\partial^2\Phi}{\partial J^2} + \frac{2\bar{\lambda}^2}{J^{5/3}} (e_r - \frac{1}{3}\text{tr}\{\mathbf{e}\}) \Psi(J) \\ & + \frac{2}{J^{5/3}} [\lambda_r^2 - \frac{1}{3}\text{tr}\mathbf{\Lambda}^2] \Psi(J) + \frac{2}{J^{5/3}} [\lambda_r^2 - \frac{1}{3}\text{tr}\mathbf{\Lambda}^2] \text{tr}\{\mathbf{e}\}J \frac{\partial\Psi}{\partial J} \\ & + [L - 3] \left\{ \frac{\partial\Psi}{\partial J} + \text{tr}\{\mathbf{e}\}J \frac{\partial^2\Psi}{\partial J^2} \right\} + \frac{\bar{\lambda}^2}{J^{2/3}} [3 + 2\text{tr}\{\mathbf{e}\} - 3 - 2\text{tr}\{\mathbf{e}\}] \frac{\partial\Psi}{\partial J}. \end{aligned} \quad (\text{A.8})$$

346 For the hydrostatic base state all the terms in square brackets in the last two lines  
347 of (A.8) vanish, and so (A.8) reduces to

$$348 \quad \sigma_1 = -p + J \frac{\partial^2\Phi}{\partial J^2} \text{tr}\{\mathbf{e}\} + \frac{2}{J} \Psi(J) (e_1 - \frac{1}{3}\text{tr}\{\mathbf{e}\}), \quad (\text{A.9})$$

349 since  $-p = \partial\Phi/\partial J$ , and  $\bar{\lambda}^2 = J^{2/3}$ .

350 The representation of the principal stress in terms of the bulk modulus  $K$  and  
 351 shear modulus  $G$  is

$$352 \quad \sigma_r = -p + K \text{tr}\{\mathbf{e}\} + G \left( \mathbf{e}_r - \frac{1}{3} \text{tr}\{\mathbf{e}\} \mathbf{1} \right), \quad (\text{A.10})$$

353 and thus we identify

$$354 \quad K = J \frac{\partial^2 \Phi(J)}{\partial J^2}, \quad G = \frac{2}{J} \Psi(J). \quad (\text{A.11})$$

### 355 References

356 Al'tshuler, L. V., Brusnikin, S. E. and Kuzmenkov, E. A., 1987. Isotherms and  
 357 Grüneisen functions for 25 metals. *J. Appl. Mech. Tech. Phys.* 28, 129-141.

358 Hill, R. 1968. On constitutive inequalities for simple materials - I., *J. Mech. Phys.*  
 359 *Solids*, 16, 229-242.

360 Keane, A., 1954. An investigation of finite strain in an isotropic material subjected  
 361 to hydrostatic pressure and its seismological applications. *Austral. J. Phys.* 7,  
 362 322-333.

363 Kennett, B.L.N., Engdahl, E.R and Buland R., 1995. Constraints on seismic  
 364 velocities in the Earth from travel times, *Geophys. J. Int.*, 122, 108–124.

365 Kennett, B.L.N. and Jackson, I., 2009. Optimal equations of state for mantle  
 366 minerals from simultaneous non-linear inversion of multiple datasets, *Phys.*  
 367 *Earth Planet. Inter.* 176, 98–108.

368 Kennett, B.L.N. and Bunge, H.-P., 2008. *Geophysical Continua*, Cambridge  
 369 University Press.

370 Poirier, J.-P. and Tarantola, A., 1998. A logarithmic equation of state, *Phys. Earth*  
 371 *Planet. Inter.* 109, 1–8.

- 372 Sakai, T., Dekura, H. and Hirao N., 2016. Experimental and theoretical thermal  
373 equations of state of MgSiO<sub>3</sub> post-perovskite at multi-megabar pressures,  
374 Scientific Reports 6, 22652. doi:10.1038/srep22652
- 375 Spencer, A.J.M., 1980 *Continuum Mechanics*, Longman.
- 376 Stacey, F.D. and Davis, P.M., 2004. High pressure equations of state with  
377 applications to the lower mantle and core, Phys. Earth Planet. Inter, 142,  
378 137–184.
- 379 Stixrude, L. and Lithgow-Bertelloni, C., 2005. Thermodynamics of mantle  
380 minerals - I. Physical Properties, Geophys. J. Int., 162, 610–632.
- 381 Vinet, P., Ferrante, J., Rose, J.H. and Smith J.R., 1987. Compressibility of Solids,  
382 J. Geophys. Res., 92, 9319-9325.

## Highlights

- New formulation of constitutive equations for deep Earth studies
- Separation of hydrostatic and deviatoric components
- Allows use of existing equations of state but with a shear modulus attached.

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