#### Highlights

Brines involved in Proterozoic U deposits (Athabasca, Canada) are targeted.

Halogens (Cl, Br, I) indicate a subaerial evaporation of seawater origin for brines.

Brines acquired excess noble gases (Ar, Kr, Xe) by interaction with clastic sediments.

Interaction with basement-derived fluid or organic matter is unlikely.

	1	Noble gases (Ar, Kr, Xe) and halogens (Cl, Br, I) in fluid inclusions from the Athabasca
1 2 3	2	Basin (Canada): Implications for unconformity-related U deposits
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#### 18 Abstract

The formation of unconformity-related uranium deposits in the Proterozoic Athabasca Basin (Canada) involved basin-scale circulation of U-bearing brines during high-grade diagenesis (150-200°C) at ~1.6-1.5 Ga. The UO<sub>2</sub> ores occur both sides of the unconformity and are associated with extensive brecciation and illite-sudoite-dravite alteration. Quartz and dolomite cementing veins and breccias are associated with alteration and mineralisation and contain a fairly uniform population of fluid inclusions characterised by variable Na:Ca and salinities of 25-35 wt.% salts and high U concentrations of up to 600 ppm U. In order to further constrain the origin of these U-rich brines, we analysed the naturally occurring isotopes of Ar, Kr and Xe, together with halogens (Cl, Br and I), K, Ca and U in irradiated quartz and dolomite samples containing representative fluid inclusions. This was achieved by the noble gas method for halogen measurement (extended <sup>40</sup>Ar-<sup>39</sup>Ar methodology) using a combination of noble gas extraction techniques.

The fluid inclusions opened by crushing quartz and dolomite samples in vacuum have similar molar Br/Cl ratios of  $5.8 \times 10^{-3}$  to  $10.4 \times 10^{-3}$ , and molar I/Cl ratios of  $1.8 \times 10^{-6}$  to  $8.2 \times 10^{-6}$ . These compositions lie over the top half of the modern-day seawater evaporation trajectory, consistent with the fluids deriving the bulk of their salinity by subaerial evaporation of seawater, beyond the point of halite saturation. The I/Cl ratios are much lower than is typical of fluids that have interacted with I-rich organic matter present in many sedimentary basins or fluid inclusions found in Mississippi Valley type (MVT) Pb-Zn ore deposits. This is significant because provided the U-rich fluid inclusions are representative of the ore-stage fluids, the low I/Cl ratios of the fluid inclusions do not favour fluid interaction with organic matter (or hydrocarbons), as a major process for localizing U mineralisation.

The majority of samples contain fluid inclusions with age-corrected  ${}^{40}Ar/{}^{36}Ar$  of between the modern atmospheric value of ~300 and 450. These values are considered

representative of the fluid's initial composition and are typical of upper crustal sedimentary formation waters. The fluid inclusions non-radiogenic <sup>84</sup>Kr/<sup>36</sup>Ar and <sup>129</sup>Xe/<sup>36</sup>Ar ratios are slightly enriched in <sup>129</sup>Xe relative to air and the fluid inclusions are estimated to contain 0.5- $17.3 \times 10^{-10}$  mol g<sup>-1</sup> <sup>36</sup>Ar which is up to twenty times the <sup>36</sup>Ar concentration of air-saturated seawater. The data are interpreted to reflect acquisition of atmospheric noble gases from sedimentary rocks and suggest acquisition of radiogenic <sup>40</sup>Ar within K-rich basement rocks, that would have been an important source of excess <sup>40</sup>Ar, was limited by temperatures of less than 200°C.

Taken together the halogen and noble gas composition of the U-bearing fluid inclusions are strongly controlled by subaerial evaporation and subsequent interaction with sedimentary rocks, showing that low temperature evaporitic brines dominated the mineralising system. Mineralisation is unlikely to have been triggered by fluid interaction with organic matter, or mixing with voluminous basement-derived fluids; however, the data do not completely preclude a role for volumetrically minor fluid or gas phases introduced by deep-seated basement faults preferentially located at the sites of mineralisation.

#### 59 Keywords

60 Unconformity-related uranium deposits

- 61 Fluid inclusions
- 62 Halogens
- 63 Noble gases
- 64 Athabasca

#### **1. Introduction**

Spectacularly high grade and large tonnage deposition of uraninite (up to 200 kt of U at 20% U on average at McArthur River) has occurred in sedimentary rocks close to the basement unconformity of the Paleoproterozoic Athabasca Basin (Saskatchewan, Canada) (Jefferson et al., 2007; Kyser and Cuney, 2008). These occurrences attest to protracted basin-scale circulation of U-bearing fluids through sedimentary and basement rocks concurrent with sediment diagenesis at temperatures of 150-200°C (e.g. Hoeve and Sibbald, 1978; Kotzer and Kyser, 1995; Derome et al., 2005; Boiron et al., 2010; Richard et al., 2012).

This style of "unconformity-related" uranium mineralisation is unique to a restricted number of Proterozoic basins including the Athabasca and Thelon basins of Canada (e.g. Renac et al., 2002; Beyer et al., 2011) and the Kombolgie Basin of Northern Australia (Derome et al., 2003a, 2007; Polito et al., 2011). However, comparable processes may have been responsible for the famous Oklo U deposits in the Franceville Basin (Gabon), which on account of their high grades and slightly greater age formed natural nuclear reactors at ~2 Ga (e.g. Shukolyukov et al., 1976; Mathieu et al., 2000; Meshik et al., 2004; Boiron et al., 2010). The occurrence of these deposits may be restricted because their genesis is facilitated by the presence of a U-rich basement (10-100 ppm U), such as underlies many Paleoproterozoic terranes (Cuney, 2010). Furthermore, high porosity, kilometre-thick quartz sandstones, that are common in the basal section of the mineralised basins, acted as essential reservoirs for voluminous saline fluids, that enabled efficient fluid convection, suitable for transport of oxidised U(VI) from basement and/or sedimentary sources to the sites of deposition (Kyser et al., 2000).

88 The  $UO_2$  ores in the Athabasca Basin are closely associated with intense illite-sudoite-89 dravite alteration in both basin and basement rocks and multiple generations of quartz and 90 dolomite that form sandstone silicifications as well as veins and breccia cements (e.g. Hoeve

and Sibbald, 1978; Pagel et al., 1980; Kotzer and Kyser, 1995; Derome et al., 2005). Similar high salinity brine inclusions have been found in sandstone silicifications, quartz and dolomite vein and breccia cements as well as secondary fluid inclusions planes in basement magmatic and metamorphic quartz from both mineralised and barren altered areas and have been implicated as the dominant fluid involved at the time of alteration and mineralisation by several workers (e.g. Pagel et al., 1980; Kotzer and Kyser, 1995; Derome et al., 2005; Mercadier et al., 2010; Richard et al., 2010, 2012).

The origin of the brines and the fluid/rock interactions they underwent have been assessed by a variety of techniques applied to fluid inclusions and their host minerals (quartz and dolomite) in a number of deposits including those studied here. Microthermometry, LA-ICP-MS and synchrotron-SRF and XANES analyses show that the brines have temperatures of 150-200°C, salinities of 25-35 wt.% salts, variable Na:Ca ratios with NaCl-rich (Na>Ca>Mg>K) and CaCl<sub>2</sub>-rich (Ca>Mg>Na>K) end-members, and that the fluid inclusions contain up to 600 ppm U(VI) (e.g. Pagel et al., 1980; Kotzer and Kyser, 1995; Derome et al., 2005; Richard et al., 2010, 2012, 2013a). Quartz and dolomite vein and breccia cements have been investigated isotopically for their O isotopic composition and these data show that quartz- and dolomite-forming fluids (trapped as fluid inclusions) are isotopically similar to the fluids responsible for deposit scale illite-sudoite-dravite alteration (Richard et al., 2013b). Furthermore, the fluid inclusions halogen (Cl, Br) composition (obtained by crush-leach ion chromatography and LA-ICP-MS analysis of individual fluid inclusions); the Cl isotope composition of fluid inclusion leachates; and the B isotope composition of dravite have been investigated. These studies favour subaerial evaporation of seawater as the dominant mechanism for generating the fluids high salinities (Richard et al., 2011; Leisen et al., 2012; Mercadier et al., 2012).

Despite these powerful constraints on the nature of the ore fluids, there remain several open questions, that will be addressed by the current study which extends the analysis of halogens in the fluid inclusions, to include I, and combines the measurements of halogens with noble gases. Firstly, it is debatable if NaCl-rich and CaCl<sub>2</sub>-rich brines represent distinct fluids, derived for example by evaporation of seawater to different degrees, or if the Ca-Na-Mg-K variation simply reflects variable fluid-rock interactions. Secondly, the basement rocks represent a potential U source (Hecht and Cuney, 2000; Mercadier et al., 2013), but the extent to which basinal fluids were convected into the basement and equilibrated with basement rocks is poorly constrained. Thirdly, the mechanism for the reduction of U(VI) to U(IV) and subsequent precipitation of UO<sub>2</sub> is still contentious. Fluid-rock interaction, fluid-organic matter interaction and mixing with deeply-sourced fluids are alternatively invoked (e.g. Hoeve and Sibbald, 1978; Wilson and Kyser, 1987; Alexandre and Kyser, 2006).

The noble gases can provide new insights on fluid interaction with basement rocks, or the presence of deeply-sourced fluid components because old K-rich basement rocks, such as the 1.8 Ga to Archean age rocks underlying the Athabasca Basin (Card et al., 2007) would be expected to have high <sup>40</sup>Ar/<sup>36</sup>Ar ratios (e.g. Kendrick et al., 2011a,b). In contrast, the relatively K-poor overlying sediments (deposited from 1.76 Ga; Ramaekers et al., 2007) that contained air saturated waters and trapped atmospheric noble gas components (e.g. Podosek et al., 1980) are more likely to preserve low <sup>40</sup>Ar/<sup>36</sup>Ar values. Iodine is an essential element for life and its incorporation into organic matter means that fluid I/Cl ratios are a sensitive indicator for fluid interaction with organic matter (Kendrick et al., 2011c). The measurement of noble gases (Ar, Xe, Kr) and halogens (Cl, Br, I) by extended <sup>40</sup>Ar/<sup>39</sup>Ar methodology provides simultaneous analysis of K, Ca and U (Kendrick, 2012), meaning noble gas and halogen signatures can be closely related to the fluids U and cation content (indicative of fluid-rock interaction). Finally, determination of Cl and Br, with I, by this methodology,

provides an independent confirmation of ion chromatographic and LA-ICP-MS Br and Cl analyses (Richard et al., 2011; Leisen et al., 2012) and additional plots that include I for assessing the extent to which the data 'fit' the suggested fluid origin by seawater evaporation.

#### 2. Geological setting, sampling and fluid inclusions

#### 2.1. The Archean to Paleoproterozoic Basement

The Athabasca Basin overlies basement rocks of the Archean to Paleoproterozoic Western Churchill Province which is further divided into the Rae Subprovince in the west and the Hearne Subprovince in the east by the Northeast-trending Snowbird tectonic zone (Hoffman, 1990; Card et al., 2007) (Fig. 1). The basement lithologies comprise Archean gneisses, Paleoproterozoic metapelites and mafic to felsic intrusions, that were all strongly metamorphosed during the ~2.0-1.9 Ga Thelon-Talston orogeny in the west, and the ~1.9-1.8 Ga Trans-Hudson orogeny in the east (Chiarenzelli et al., 1998; Annesley et al., 2005; Card et al., 2007). Most of the unconformity-related uranium deposits in the Athabasca Basin are located in the vicinity of the Wollaston-Mudjatik transition zone (WMTZ) that separates the Wollaston and Mudjatik lithostructural domains of the Hearne Subprovince (Annesley et al., 2005; Fig. 1).

#### 2.2. The Paleoproterozoic Athabasca Basin

The sedimentary rocks of the Athabasca Basin, known as the Athabasca Group, unconformably overly the crystalline basement, and sediment deposition started at ~1.75 Ga (Ramaekers et al., 2007; Fig. 1). The current maximum thickness of the sedimentary cover is ~1.5 km (Rumple Lake), but it is thought to have reached a maximum thickness of ~5 km, based on P-T estimations from fluid inclusion studies (Pagel, 1975; Derome et al., 2005). The oldest sediments in the Athabasca Group comprise a ~1 km thick basal sequence of fluvial to

marginal marine quartz-rich sandstones known as the Fair Point, Read, Smart and Manitou Falls Formations. The overlying Lazenby Lake and Wolverine Point Formations comprise marine sandstones, phosphatic siltstones and phosphatic mudstones. These are overlain by fluviatile sandstones of the Locker Lake and Otherside Formations, shales of the Douglas Formation and finally stromatolitic carbonates of the Carswell Formation (Ramaekers et al., 2007).

#### 2.3. The Mesoproterozoic unconformity-related uranium deposits in the Athabasca Basin

The peculiarity of unconformity-related uranium deposits in the Athabasca Basin is that they can reach extremely high grade and tonnage. For example, the giant McArthur River deposit in the Athabasca Basin, which is included in the current study, has produced ~200,000 tonnes of U ore at a grade of 20 wt.% U (Jefferson et al., 2007).

The Athabasca unconformity-related U ores are typically located within ~100-400 m of the unconformity between the Archean to Paleoproterozoic basement and the Paleoproterozoic sedimentary cover, close to graphite-rich basement faults that occur preferentially within metapelite units. Of the deposits included in this study, the P-Patch and Eagle Point deposits are entirely hosted by crystalline basement (Mercadier et al., 2010; Cloutier et al., 2011). The McArthur River deposit occurs on both sides of the unconformity (e.g. Derome et al., 2005). The Rabbit Lake is located at the very margin of the preserved sedimentary cover and is presently entirely hosted in the basement but may have been spatially connected to the unconformity prior to erosion (e.g. Hoeve and Sibbald, 1978; Heine, 1986; Alexandre et al., 2005).

The main alteration features associated with  $UO_2$ -ores include: (i) partial to complete replacement of the initial basement minerals (K-feldspar, biotite, plagioclase, hornblende) by an illite + sudoite (Mg-chlorite) ± dravite (Mg-tourmaline) assemblage (e.g. Alexandre et al.,

2005; Mercadier et al., 2010), as well as the precipitation of quartz, illite, sudoite and dravite in the sandstones (e.g. Kotzer and Kyser 1995; Lorilleux et al., 2002; Kister et al., 2006); (ii) quartz dissolution in basement rocks as well as in the sandstones with hydraulic and chemical brecciation (e.g. Lorilleux et al., 2003; Mercadier et al., 2010); (iii) precipitation of UO<sub>2</sub> in the alteration-related porosity as well as in veins and breccias also cemented by dravite  $\pm$  quartz  $\pm$ dolomite  $\pm$  bitumen (e.g. Hoeve and Sibbald, 1978; Derome et al., 2005) and (v) rare earth element mobilization and redistribution among aluminium phosphate-sulphate minerals such as florencite, goyazite and svanbergite and UO<sub>2</sub> (e.g. Fayek and Kyser, 1997; Gaboreau et al., 2007; Mercadier et al., 2011a). A simplified mineral paragenetic sequence summarizing petrographical observations of U deposits throughout the Athabasca Basin is shown in Fig. 1.

A wide range of ages spanning 1.6-0.9 Ga have been reported for U-related mineralisation and alteration minerals in the Athabasca Basin (e.g. Cummings and Krstic, 1992; Fayek et al., 2002; Alexandre et al., 2009). However, the most recent geochronological data show that pre-ore illite and sudoite preserve maximum  ${}^{40}Ar/{}^{39}Ar$  ages of ~1670 Ma and syn-ore illite preserves a maximum  ${}^{40}\text{Ar}/{}^{39}\text{Ar}$  age of ~1580 Ma that is similar to the maximum U-Pb ages of  $1540 \pm 20$  Ma obtained for uraninite (Alexandre et al., 2009). These data can be interpreted to indicate alteration and mineralisation spanned a period of <150 Ma between 1670 and 1500 Ma, and the younger ages recorded by both the  ${}^{40}\text{Ar}/{}^{39}\text{Ar}$  and U/Pb systems result from remobilisation of UO<sub>2</sub> and partial <sup>40</sup>Ar or Pb loss, which has probably occurred multiple times up until the present day (e.g. Alexandre et al., 2009; Fayek et al., 2002; Mercadier et al., 2011b).

#### 2.4. Quartz and dolomite cements and their fluid inclusions in the ore-forming process

Many workers agree quartz cements are closely spatially associated with the main alteration minerals (illite, sudoite, dravite) and UO<sub>2</sub> (e.g. Dahlkamp, 1978; Hoeve and

Sibbald, 1978; Wilson and Kyser, 1987; Kotzer and Kyser, 1995; Fayek and Kyser, 1997; Kyser et al., 2000; Lorilleux et al., 2002; Derome et al., 2005; Kyser and Cuney, 2008; Ng et al., 2013). Furthermore, carbonates (including dolomite) have been consistently described as gangue minerals in the Rabbit Lake, Shea Creek and McArthur River deposits (Hoeve and Sibbald, 1978; Pagel et al., 1980; Hoeve et al., 1986; Heine, 1986; Lorilleux et al., 2002; Derome et al., 2005). As a consequence of this close association, fluid inclusions in the quartz and dolomite cements have been targeted as representative of the dominant fluid involved at the time of alteration and mineralisation. The following observations indicate that NaCl-rich and CaCl<sub>2</sub>-rich brines were the dominant fluid involved from the pre-ore stage to the syn-ore stage and the post-ore stage: (i) sandstone silicifications, which are unanimously recognized as preceding the main UO<sub>2</sub> stage show primary NaCl-rich and CaCl<sub>2</sub>-rich brine inclusions at Cluff Lake and McArthur River deposits (Pagel, 1975; Kotzer and Kyser, 1995; Derome et al., 2005); (ii) the increasing intensity of illite-sudoite-dravite alteration is related to the increasing abundance of secondary NaCl-rich and CaCl<sub>2</sub>-rich brine inclusion planes in the magmatic and metamorphic quartz from basement rocks at the P-Patch deposit (Mercadier et al., 2011); (iii) quartz and dolomite cements in barren and mineralized veins and breccias are all filled by primary and pseudosecondary NaCl-rich and CaCl<sub>2</sub>-rich brine inclusions in all the deposits studied here (Pagel et al., 1980; Kotzer and Kyser, 1995; Derome et al., 2005; Richard et al., 2011); (iv) LA-ICP-MS and synchrotron-SRF and XANES analyses have been applied to those NaCl-rich and CaCl<sub>2</sub>-rich brine inclusions in both barren and mineralised samples and showed that they contain up to 600 ppm U(VI) (Richard et al., 2010, 2012, 2013a); (v) the O isotopic composition and these quartz and dolomite cements shows that cement-forming fluids (i.e. NaCl-rich and CaCl<sub>2</sub>-rich brines) are isotopically equilibrated with illite, sudoite and dravite (Richard et al., 2013b) and (v) the B isotopic composition of dravite as well as Cl/Br ratios coupled with Cl isotope compositions of NaCl-rich and CaCl<sub>2</sub>-rich brine inclusions in quartz and dolomite cements indicate a common evaporated seawater origin for the dravite-forming fluids and the NaCl-rich and CaCl<sub>2</sub>-rich brines (Richard et al., 2011; Leisen et al., 2012; Mercadier et al., 2012).

Some other workers have concluded from crosscutting relationships that quartz and dolomite cements postdate the main UO<sub>2</sub> stage meaning their fluid inclusions cannot be used to identify the ore fluids (e.g. Alexandre et al., 2005; Cloutier et al. 2009). Indeed, the relative timing of the alteration minerals, quartz and dolomite cements and the  $UO_2$  can be ambiguous because mineral deposition is often cyclic (Dahlkamp, 1978; Hoeve and Sibbald, 1978) and the different stages are variably expressed in basement and sandstone lithologies (e.g. Kyser and Cuney, 2008; Alexandre et al., 2009). In addition, recurrent episodes of quartz dissolution have been identified (e.g. Derome et al., 2005) meaning the fluid inclusions in quartz and dolomite may not have recorded the entire fluid story. As a consequence, it is always debatable whether a given quartz or dolomite sample contains fluid inclusions which are fully representative of the fluids involved at the main UO<sub>2</sub> stage. However, as mentioned above, all fluid inclusion studies point to NaCl-rich and CaCl2-rich brines as the dominant fluid involved.

#### 2.5. Sampling for the current study

Seven samples from McArthur River (MAC54Qz), P-Patch (P48-2, P48-5), Rabbit Lake (RBL2 and RBL7) and Eagle Point (ES287-8, H3042-1), which are 4 of the best characterised deposits in the Athabasca Basin, were selected for the current study (Fig. 1; Fig. 2; Table 1). These samples show typical quartz and dolomite cements in veins and breccias and typical NaCl-rich and CaCl<sub>2</sub>-rich brine inclusions as described in details elsewhere within alteration haloes as well as in barren and mineralized samples (e.g. Pagel et al., 1980; Kotzer and Kyser, 1995; Derome et al., 2005; Richard et al., 2010, 2011; Mercadier et al., 2011).

The samples were selected from varying depths into the basement lithologies, to a maximum of 263 m below the unconformity (H3042-1). The samples show variable degree of alteration, from relatively unaltered (H3042-1) to highly altered "bleached" (illite-sudoite-dravite alteration) (ES-287-8, P-48-5). This was undertaken to evaluate if a basement noble gas signature could be detected. Some samples contain quartz cement without dolomite (ES287-8, H3042-1, P48-2, P48-5, MAC54Qz) while others show quartz and dolomite cements, dolomite postdating quartz (RBL2 and RBL7). Quartz and dolomite cements in both veins (ES287-8, H3042-1, P48-2, P48-5) and breccias (MAC54Qz, RBL2 and RBL7) were selected. Together with quartz, veins and breccias are cemented by dravite or dravite+sulphides in samples H3042-1 and MAC54Qz respectively. Textural evidence suggests that dravite predates quartz which in turn predates sulphides. Despite the close association of the selected samples with mineralisation, none of the selected samples contain UO<sub>2</sub> ore within the hand-specimen sized section available from drill core. Barren samples were specifically selected to avoid possible artefacts related to post-entrapment production of nucleogenic noble gases modifying the fluid inclusions <sup>36</sup>Ar concentration (See Section 3).

The fluid inclusion characteristics of the selected samples have been described in Richard et al. (2011) and are summarised below. The selected quartz and dolomite samples are dominated by primary and pseudosecondary fluid inclusions of 5 to 25  $\mu$ m in size in which a vapour phase occupies ~10% of the inclusion volume. The majority of inclusions have two phases (liquid + vapour) but up to ~25% contain three phases (liquid + vapor + halite) (Fig. 2; Table 1). A few of the inclusions contain haematite and/or phyllosilicate minerals (illite and chlorite as identified by Raman spectrometry by Derome et al., 2005) that are also interpreted as daughter minerals due to their constant volume within the fluid inclusions. This is significant because, it demonstrates that some of the fluid inclusions within the selected samples are saturated with respect to alteration minerals. The fluid inclusions

have a restricted range of salinity of ~25-35 wt.% salts, and homogenise into the liquid phase generally between 90 and 170°C. The relative abundances of different fluid inclusion types representative of NaCl and CaCl<sub>2</sub> brines, together with average salinities of fluid inclusions in each sample, are summarised in Table 1. In addition to the primary fluid inclusions that are of interest for investigating U mineralisation, the samples contain a low abundance of lowsalinity secondary fluid inclusions similar to those described by Derome et al., (2005). The low abundance (<1%) and low salinity (1-10 wt.% eq. NaCl) of these secondary fluid inclusions implies they will make a negligible contribution to the Cl, Br, I, K, Ca and U compositions measured by bulk analysis of the fluid inclusion bearing samples in this study. However, each sample was analysed sequentially, by multiple in vacuo crushing analyses to test for sample heterogeneity and help asses if the secondary fluid inclusions significantly influenced the noble gas signatures (see Section 3).

#### 3. Methodology

The halogens (Cl, Br, I), Ca, K and U were measured simultaneously with the naturally occurring noble gas isotopes of Ar, Kr and Xe using the noble gas method which is described in detail by Kendrick (2012) (see also Böhlke and Irwin (1992a) and Johnson et al. (2000)). Fluid inclusion bearing quartz and dolomite samples of 30-110 mg were irradiated for 50 hours in position 5c of the McMaster Nuclear Reactor, Canada (irradiation designated UM#37 on the 5<sup>th</sup> January 2010), to produce noble gas proxy isotopes ( ${}^{38}Ar_{Cl}$ ,  ${}^{80}Kr_{Br}$ ,  ${}^{128}Xe_{I}$ ,  ${}^{37}Ar_{Ca}$ ,  ${}^{39}Ar_{K}$  and  ${}^{134}Xe_{U}$ ) that are measured by noble gas mass spectrometry (Kendrick, 2012).

The irradiated samples were loaded into modified Nupro® valves, baked under ultrahigh vacuum at ~100 °C for 24-48 hrs to remove loosely adsorbed noble gases, and then analysed by in vacuum crushing. Short crushing experiments in Nupro® valves release naturally occurring and neutron-produced noble gas isotopes from fluid inclusions and contributions from the mineral matrix are negligible (Kendrick and Phillips, 2009). Each sample was analysed by multiple crushes, enabling compositional variations resulting from fluid inclusion heterogeneity to be assessed (Section 2.5). The crushed residues of 6 quartz samples were then loaded into Sn-foil packets and fused in a 20 minute heating step at 1600°C using a tantalum resistance furnace. Fusion removes noble gases from remaining fluid inclusions and the mineral matrix, meaning the combined *in vacuo* crushing and fusion analyses provide complementary information.

The noble gas isotopes released by *in vacuo* crushing and fusion were purified by standard gettering procedures over a period of 20 minutes and measured in peak jumping mode on the MAP-215 noble gas mass spectrometer at the University of Melbourne. The isotopes of Ar where measured on a Faraday detector whereas the less abundant isotopes of Kr and Xe were measured on the more sensitive electron multiplier detector (Kendrick, 2012).

Ratios of Cl to Br, I, Ca, K and U are obtained from the measured noble gas proxy isotopes and monitored production ratios. Production ratios for <sup>38</sup>Ar<sub>Cl</sub>/Cl, <sup>37</sup>Ar<sub>Ca</sub>/Ca and <sup>39</sup>Ar<sub>K</sub>/K were monitored using the Hb3Gr and GA1550 <sup>40</sup>Ar-<sup>39</sup>Ar flux monitors, whereas the <sup>128</sup>Xe<sub>I</sub>/I and <sup>80</sup>Kr<sub>Br</sub>/Br production ratios were monitored using four scapolite standards (Kendrick, 2012; Kendrick et al., 2013). Standard corrections were made for Ar interference reactions based on K-glass and CaF<sub>2</sub> monitors and U-derived <sup>84</sup>Kr (Kendrick, 2012). Unfortunately, fusion analyses were undertaken more than a year after sample irradiation, meaning <sup>37</sup>Ar, which has a half life of 35 days could not be measured. Fluid inclusion <sup>36</sup>Ar, <sup>40</sup>Ar and U concentrations have been calculated from the measured ratios (X/Cl) and the fluid inclusion salinity determined via microthermometry and given in Table 1.

The noble gas method provides very low detection limits, and machine blanks for crushing samples in modified Nupro® valves that typically account for <1% of sample gas. The Br/Cl and I/Cl ratios are reported relative to scapolite standards that have

reproducibilities of  $\sim 3\%$  for Br/Cl and  $\sim 5\%$  for I/Cl (1 $\sigma$ ; see Kendrick 2012; Kendrick et al., 2013). The precision of the K/Cl, Ca/Cl and U/Cl measurements were more variable and were limited by the low abundance of the noble gas proxy isotopes  $({}^{39}Ar_{K}, {}^{37}Ar_{Ca} \text{ and } {}^{134}Xe_{U})$ available for analysis in several cases. The U measurements are based on a  $^{134}Xe_{II}/U$ production ratio calculated from the monitored neutron fluence and U cross section, rather than a U monitor (Kendrick 2012). This means that while real differences in the U concentration can be precisely monitored, the absolute U values could be systematically out by up to  $\sim 20\%$  in all the samples. The accuracy of the fluid inclusion concentrations of U,  $^{40}$ Ar and  $^{36}$ Ar is further limited to the ~30% level by the variability of fluid inclusion salinities within each sample (~25-35 wt.% salts; see Section 2.5).

Finally, post entrapment modification of fluid inclusion Ar-isotope signatures are considered to be a minor artefact in this study. Firstly, the low K content of the samples means that despite the samples great age (Section 2.3), post-entrapment production of radiogenic  $^{40}$ Ar<sub>R</sub> is negligible, necessitating a correction of less than 5% for the majority of samples (Table 2). Secondly, we present data for samples situated >1 m from U ore, which minimises the possible influence of post-entrapment nucleogenic reactions modifying fluid inclusion <sup>36</sup>Ar concentrations, because neutrons have an attenuation distance of ~1 m in silicate rocks (Ballentine and Burnard, 2002; Hu et al., 2009).

#### **4. Results**

The halogen and noble gas results obtained by *in vacuo* crushing of quartz and dolomite hosted fluid inclusions are summarised in Table 2, and the results obtained by fusing selected crushed quartz samples are summarised in Table 3.

Quartz and dolomite hosted fluid inclusions extracted by *in vacuo* crushing have molar Br/Cl values between  $(5.8 \pm 0.1) \times 10^{-3}$  and  $(10.4 \pm 0.1) \times 10^{-3}$ , I/Cl values between  $(1.8 \pm 0.1) \times 10^{-3}$   $0.1) \times 10^{-6}$  and  $(8.2 \pm 0.2) \times 10^{-6}$  (Fig. 3), and K/Cl values between  $0.017 \pm 0.005$  and  $0.056 \pm 0.002$  that are typical of crustal fluids (Fig. 4; Fig. 5; Fig. 6). Quartz-hosted fluid inclusions extracted by *in vacuo* crushing have Ca/Cl values of between  $0.06 \pm 0.04$  and  $0.19 \pm 0.06$ (Fig. 4; Fig. 5; Fig. 6), but Ca data are not reported for dolomite samples for which there was a significant matrix contribution. The *in vacuo* crushing data show fluid inclusion U/Cl values vary from  $2.5 \times 10^{-6}$  to  $1.4 \times 10^{-4}$  in the majority of samples (Fig. 5). Based on the mean fluid inclusion salinities given in Table 1, the fluid inclusions are therefore indicated to contain between 3 and 170 ppm U.

The *in vacuo* crushing analyses yield imprecise  ${}^{40}$ Ar- ${}^{39}$ Ar ages of up to 4 Ga, that are much older than the age range reported for U-related alteration minerals (Cummings and Krstic, 1992; Fayek et al., 2002; Alexandre et al., 2009), demonstrating that excess  ${}^{40}$ Ar is trapped in the fluid inclusions [excess  ${}^{40}$ Ar ( ${}^{40}$ Ar<sub>E</sub>) is defined as parentless non-atmospheric  ${}^{40}$ Ar that was not generated by decay of  ${}^{40}$ K within the sample]. The  ${}^{40}$ Ar/ ${}^{36}$ Ar ratios obtained by *in vacuo* crushing range from 302 to 480 and are higher than the modern atmospheric value of 296 and considered fairly representative of the fluid inclusions initial composition at the time of trapping (Table 2). The maximum correction possible for post-entrapment production of  ${}^{40}$ Ar within the fluid inclusions is only ~5%, based on the measured K abundances of the fluid inclusions and an assumed age of 1.6 Ga (Table 2).

In vacuo crushing yields fluid inclusion  $Cl^{36}Ar$  values of between  $(2.7 \pm 0.1) \times 10^{6}$ and  $(86 \pm 1) \times 10^{6}$  and repeat crushing analyses of the dolomite and quartz samples show  $^{40}Ar^{36}Ar$  and  $Cl^{36}Ar$  are correlated in several of the samples (Fig. 6). The majority of the Ar mixing lines converge on an end-member with low  $Cl^{36}Ar$  and  $^{40}Ar^{36}Ar$  of close to the modern atmospheric value of 296, and the slopes of Ar mixing lines forced through the atmospheric  $^{40}Ar^{36}Ar$  ratio represent the samples average  $^{40}Ar_{E}/Cl$  ratio (Fig. 6).

The highest <sup>40</sup>Ar/<sup>36</sup>Ar values and Cl/<sup>36</sup>Ar values obtained by *in vacuo* crushing each sample is suggested as most representative of the NaCl and CaCl<sub>2</sub> fluid inclusions (Fig. 6). The sample specific  $Cl/^{36}Ar$  versus  $^{40}Ar/^{36}Ar$  mixing lines (Fig. 6), probably result from mixing the dominant NaCl and CaCl<sub>2</sub> brine inclusions with low salinity secondary fluid inclusions present in the samples (see Section 2.5); however, the mixing lines could also result if a modern air contaminant was trapped in voids within the samples. Based on combining the fluid inclusions  $^{40}\text{Ar}_{\text{E}}/\text{Cl}$  ratios of between 0.04  $\times$  10  $^{-6}$  and (2.7  $\pm$  0.3)  $\times$  10  $^{-6}$ (Table 2) with the mean salinity of brine inclusions in each sample, we calculate the fluid inclusions contain  $0.02\times 10^{\text{-8}}$  to  $1.4\times 10^{\text{-8}}$  mol g  $^{\text{-1}}$  of  $^{40}\text{Ar}_{\text{E}}.$  Based on the maximum Cl/ $^{36}\text{Ar}$ value obtained for each sample (considered representative of the NaCl and CaCl<sub>2</sub> brine inclusions) and the mean fluid inclusion salinity, we obtain fluid inclusion <sup>36</sup>Ar concentrations of  $0.5\times10^{\text{--}10}$  to  $17.3\times10^{\text{--}10}$  mol g^{\text{--}1} (Table 2). 

The fluid inclusion <sup>129</sup>Xe/<sup>36</sup>Ar and <sup>84</sup>Kr/<sup>36</sup>Ar values obtained by *in vacuo* crushing are reported as fractionation values relative to air (Fig. 6). The fractionation values are defined as  $F(X/^{36}Ar)_{air} = (X/^{36}Ar)_{sample}/(X/^{36}Ar)_{air}$ , where  $X = ^{84}Kr$  or <sup>130</sup>Xe, and air has  $F(^{129}Xe/^{36}Ar)_{air}$ and  $F(^{84}Kr/^{36}Ar)_{air}$  values of one. The fluid inclusion  $F(^{84}Kr/^{36}Ar)_{air}$  ratios are close to 1 but  $F(^{129}Xe/^{36}Ar)_{air}$  varies from ~1 to 2.7 ± 0.2 demonstrating a significant enrichment in <sup>129</sup>Xe relative to air (Table 2).

Fusing previously crushed samples at 1600°C released slightly less <sup>38</sup>Ar<sub>Cl</sub> than *in vacuo* crushing and similar proportions of <sup>36</sup>Ar as *in vacuo* crushing (Table 2; Table 3). The fusion data are characterised by Br/Cl ratios that are typically 30-40 % lower than the *in vacuo* crushing values and K/Cl and U/Cl values that are more variable than the *in vacuo* crushing values (Table 2; Table 3). The maximum K/Cl ratio of 6.5 obtained by fusion (Table 3) is higher than can be explained by decrepitation of fluid inclusions which have Na- or Carich compositions (Richard et al., 2010) and are indicated to have K/Cl of <0.1 by the *in* 

*vacuo* crushing data (note that pure sylvite has K/Cl of 1). Furthermore, fusion yields 415 imprecise  ${}^{40}$ Ar- ${}^{39}$ Ar ages of 0.8-1.4 Ga that overlap the range of ages of 1.2-1.6 that have 416 previously been reported for K-bearing alteration minerals (Alexandre et al., 2009), showing 417 mineral impurities make a significant contribution to the irradiation produced noble gases 418 released by fusion (Table 3). Finally,  $F({}^{84}$ Kr/ ${}^{36}$ Ar)<sub>air</sub> and  $F({}^{129}$ Xe/ ${}^{36}$ Ar)<sub>air</sub> ratios obtained by 419 fusion are uniformly greater than air with the maximum  $F({}^{129}$ Xe/ ${}^{36}$ Ar)<sub>air</sub> of 52 demonstrating a 420 very significant enrichment of atmospheric Xe relative to Ar (Table 3).

**5. Discussion** 

#### 5.1. Comparison of in vacuo crushing and fusion data

5.1.1 Fluid inclusion  $^{40}$ Ar/ $^{36}$ Ar ratios and  $^{40}$ Ar- $^{39}$ Ar ages

The simplest interpretation of the argon data is that the age-corrected <sup>40</sup>Ar/<sup>36</sup>Ar ratios of <450 and the K/Cl ratios of 0.02-0.06 obtained by *in vacuo* crushing are representative of the fluid inclusions at the time of trapping (Table 2). In contrast, fusing previously crushed samples gave a maximum K/Cl of 6.5, and U/Cl ratios of 200-300 times higher than *in vacuo* crushing, suggesting <sup>39</sup>Ar and <sup>134</sup>Xe were released from K and U mineral impurities in the crushed samples by heating (e.g. Kendrick et al., 2001; 2006). The similarity of apparent ages of 0.8-1.4 Ga obtained by fusion of the quartz samples and other K-bearing minerals (Alexandre et al., 2009) is consistent with the ages having been obtained from mineral impurities within the quartz rather than the fluid inclusions. Given the mineral impurities must be very small to have avoided detection during careful sample preparation under a binocular microscope, and small mineral impurities have very low retentivities to <sup>40</sup>Ar, the apparent ages are likely to represent cooling ages and are not related to the timing of mineralisation (see Kendrick et al., 2006). One possible complexity in the above interpretation is that the  ${}^{40}$ Ar/ ${}^{36}$ Ar measured by *in vacuo* crushing might be significantly higher than the fluid inclusions initial value if significant  ${}^{40}$ Ar was lost from phyllosilicate daughter minerals into the host fluid inclusion. However, as crustal  ${}^{40}$ Ar/ ${}^{36}$ Ar ratios vary from ~300 up to at least 50,000 (Kendrick et al., 2011b) this does not affect interpretation of the current data that have measured  ${}^{40}$ Ar/ ${}^{36}$ Ar ratios of <480 and preferred values of <460 after a correction for 1.6 Ga of post entrapment in growth (Table 2).

#### 5.1.2. Halogen ratios

The Br/Cl values obtained by *in vacuo* crushing are 30-40 % higher than the values obtained by fusion (Table 2; Table 3). It is difficult to establish the cause of this discrepancy with certainty; however, our favoured explanation is that *in vacuo* crushing releases noble gases from fluid inclusions and fusion releases noble gases from fluid inclusions in the crushed sample and mineral impurities in the quartz matrix, phyllosilicates or mica, which must explain K/Cl ratios of >1 measured by fusion (Table 3), and could also contribute significant Cl (200-500 ppm Cl in illite and sudoite the after Richard et al., 2011), with very little Br (e.g. Kendrick, 2012), that could potentially reduce the Br/Cl ratio measured by fusion.

An alternative explanation that high Br/Cl ratios could result from retention of <sup>38</sup>Ar<sub>Cl</sub> in halite daughter minerals during *in vacuo* crushing (Kendrick et al., 2001) is possible, but:(i) samples containing fluid inclusions with daughter minerals sometimes give the same Br/Cl ratios by both techniques (Fisher and Kendrick, 2008), and *in vacuo* crushing sometimes gives higher Br/Cl than stepped heating samples hosting fluid inclusions without daughter minerals (Kendrick, unpublished data), suggesting daughter crystals are minor <sup>38</sup>ArCl reservoirs; (ii) daughter mineral-bearing inclusions are minor in some of the current samples and were not observed in samples ES287-10, RBL2Qz or RBL2Carb (Table 1; Section 2.5); and (iii)
whereas Br/Cl measured by stepped heating is systematically lower than the *in vacuo* crushing
data, the I/Cl ratio obtained by fusing ES287-10 and H304-2 is higher than obtained by *in vacuo* crushing (Table 2; Table 3). This difference cannot be explained by the retention of
<sup>38</sup>ArCl in halite daughter minerals, but could potentially be explained by the presence of
organic contaminants and mineral impurities in the quartz matrix.

A final possibility is that fluid inclusions with different sizes, or different occurrences (e.g. isolated versus annealed fractures), have different Br/Cl and I/Cl ratios, and that *in vacuo* crushing and fusion release noble gases from these fluid inclusions differently. While we cannot yet preclude this possibility, if it were true we would expect non-systematic variation in Br/Cl between the two techniques, whereas to date, every time there has been a discrepancy the lower Br/Cl ratios is obtained by stepped heating (this study; Kendrick et al., 2001; 2007).

In summary, we believe the *in vacuo* crushing data emphasised throughout the remainder of this discussion are fairly representative of the fluid inclusions in our samples, and that Cl-bearing mineral impurities in the quartz matrix probably contribute to the fusion results. For simplicity we focus on the *in vacuo* crushing results; however, it should be noted that this does not greatly change the interpretations below and sample average Br/Cl and I/Cl, obtained by combining the *in vacuo* crushing and fusion data in Table 2 and Table 3, would be within ~20 % of the *in vacuo* crushing analyses.

5.2. Origin of fluid inclusion salinity

The fluid inclusion molar Br/Cl versus I/Cl values determined for quartz- and dolomite-hosted brine inclusions by *in vacuo* crushing using the noble gas method are compared with modern seawater evaporation trends and the compositions of other important

fluid and mineral reservoirs in Fig. 3. The fluid inclusion data lie over the top part of the
modern seawater evaporation trajectory (Zherebtsova and Volkova, 1966).

The new Br/Cl values obtained by *in vacuo* crushing  $(5.8 \times 10^{-3} \text{ to } 10.4 \times 10^{-3})$  are higher than those measured previously on the same samples by crush-leach ion chromatography (3  $\times$  10<sup>-3</sup> to 6  $\times$  10<sup>-3</sup>; Richard et al., 2011; Fig. 3), or determined for individual fluid inclusions in from McArthur River by LA-ICP-MS (1  $\times$  10<sup>-3</sup> to 5  $\times$  10<sup>-3</sup>; Leisen et al., 2012); however, crush-leach ion chromatography has given Br/Cl of up to  $10 \times$  $10^{-3}$  for Eagle Point dolomite and  $8 \times 10^{-3}$  for McArthur River dolomite (Richard et al., 2011). Therefore although the Br/Cl ratios obtained in this study are higher than those reported previously (Fig. 3), all of the techniques give overlapping results on large sample sets, and some of the difference for the common samples analysed by both techniques could be related to sample heterogeneity. Further work to inter-calibrate these different techniques for measuring Br and Cl (e.g. Hammerli et al., 2013; Kendrick et al., 2013); and detailed comparison of the methods used for extracting fluid inclusions is clearly desirable (e.g. crushleach; in vacuo crushing; stepped heating). However, as all the techniques give Br/Cl of mostly greater than seawater, the differences do not strongly influence the interpretation of the data in the current study. We focus on the in vacuo crushing Br/Cl data obtained here because this technique has very high internal precision and provides simultaneous I/Cl analyses.

The *in vacuo* crushing Br/Cl and I/Cl ratios lie over the top part of the modern day seawater evaporation trajectory (Fig. 3). Taken together with the high chlorinity of the fluid inclusions (~5000 mmol/kg solution in Fig. 3; ~25-35 wt.% salts) this provides strong evidence that the fluids derived the bulk of their salinity by subaerial evaporation of seawater beyond the point of halite saturation (Fig. 3). However, despite the indicated importance of seawater evaporation, this mechanism produces residual brines with maximum salinities of

512 ~30 wt.% salts (Zherebtsova and Volkova, 1966; Nissenbaum, 1977). The existence of halite 513 saturated fluid inclusions with visible daughter crystals and total salinities of ~35 wt.% salts 514 in these samples (Fig. 2), therefore requires a further mechanism to increase the fluids salinity 515 above the point of halite saturation.

The most likely mechanism for increasing the brines salinity above ~30 wt.% salts was probably H<sub>2</sub>O consumption during illitisation and chloritisation of feldspar during basement alteration. Fluid inclusions show progressively higher salinities in the areas of the basement that have been most intensely altered by illitisation and chloritisation (Mercadier et al., 2010). Another mechanism that could have increased the salinity of the evaporitic brine above 30 wt.% would have been the dissolution of evaporites that might have been present within overlying sediments. However, although dissolution of evaporites would increase the fluids salinity, it would also move the fluids Br/Cl and I/Cl composition back down the seawater evaporation trajectory (Fig. 3). As a result, evaporite dissolution would lead to the fluids with the highest salinities tending to have the lowest Br/Cl values, which is not observed (Fig. 3).

The measurement of I/Cl, in addition to Br/Cl, is important because sedimentary formation waters can evolve high I/Cl and Br/Cl values as a result of interaction with organic matter (Kendrick et al., 2011c). The low I/Cl values of the U-related fluids indicate very little interaction with organic-rich sediments or hydrocarbons and the elevated Br/Cl cannot be attributed to an organic Br component as previously suggested by Pagel and Jaffrezic (1977) and Pagel et al. (1980). The lack of organic I and Br components in the ore fluids related to unconformity U mineralisation is in stark contrast to the majority of formation waters and ore fluids related to sediment-hosted Zn-Pb mineralisation in Mississippi Valley-type ore deposits which have I/Cl of 10-1000  $\times$  10<sup>-6</sup> (Fig. 3; Kendrick et al., 2011c). The new results demonstrate the NaCl-rich and CaCl<sub>2</sub>-rich brines do not contain a significant organic component.

In summary, the halogen data indicate seawater evaporation as the dominant source of salinity and that illitisation and chloritisation were probably important for further increasing fluids salinity. The dominant source of salinity is therefore similar to that suggested by the lower Br/Cl ratios obtained by ion chromatography methods previously (Richard et al., 2011). The main uncertainty resulting from the different Br/Cl ratios (Fig. 3) is the extent to which minor processes such as illitisation or evaporite dissolution have modified the bittern brines.

#### 5.3. Origin of fluid inclusion cation (Ca, K, U) composition

The bulk fluid inclusion molar Ca/Cl and K/Cl values determined by *in vacuo* crushing of quartz and dolomite samples are shown in Fig. 4 together with the approximate fields of the most NaCl-rich and CaCl<sub>2</sub>-rich fluid inclusion end-members analysed by LA-ICP-MS (Richard et al., 2010) and microthermometry (Derome et al., 2005).

Most of the *in vacuo* crushing analyses have Ca/Cl of less than 0.15 that suggest the samples are dominated by NaCl-rich fluid inclusions even though CaCl<sub>2</sub>-rich inclusions have been observed in some of the samples (Table 1). The extreme NaCl-rich and CaCl<sub>2</sub>-rich end-member compositions can only be defined by analysing individual fluid inclusions (e.g. Richard et al., 2010). However, the *in vacuo* crushing data do not lie on a simple mixing trend between the suggested end-members (Fig. 4), which is most easily explained if the K/Cl composition of NaCl-rich fluid inclusions is more variable than previously suggested (Fig. 4).

The fluid inclusion molar Ca/Cl and K/Cl values are plotted against Br/Cl, I/Cl and U/Cl values in Fig. 5. Evaporation trends of MgSO<sub>4</sub>-rich seawater (similar to modern day seawater) and CaCl<sub>2</sub>-rich seawater are shown in Fig. 5, because the actual composition of seawater at the time of U mineralisation is unknown and seawater is suggested to have oscillated between these types over Earth history (Lowenstein et al., 2001; Hardie 2003; Kovalevych et al., 2006). In contrast, to Br/Cl and I/Cl data shown in Fig. 3, the fluid

inclusions Ca/Cl and K/Cl values cannot be easily related to any particular stage of either seawater evaporation trajectory (Fig. 5). Both the K/Cl and Ca/Cl composition of the fluid inclusions could be influenced by seawater evaporation; however, intensive modification of K/Cl and Ca/Cl values caused by fluid-rock interactions during diagenesis is likely to account for large variation of K/Cl and Ca/Cl values. Illitisation of basin and basement rocks has been documented in the Athabasca Basin and is probably a major cause of K/Cl variation (Kister et al., 2006; Mercadier et al., 2010). Variations in formation water Ca/Cl are often attributed to either dolomitisation of limestones or albitisation of detrital plagioclase and basement rocks (Davisson and Criss, 1996; Houston et al., 2011). Detrital plagioclase are nearly absent from the preserved sedimentary succession (Ramaekers et al., 2007) and no albitisation of basement plagioclase is observed (Mercadier et al., 2010). However, stromatolitic to massive dolostones are described in the Carswell Formation (Ramaekers et al., 2007).

The range of U concentrations determined for fluid inclusions by *in vacuo* crushing in this study (3-169 ppm) is within the range of U concentrations (0.2-600 ppm) measured for individual inclusions by LA-ICP-MS (Richard et al., 2010, 2012). The U/Cl value is not correlated with either Ca/Cl or K/Cl (Fig. 5) or any other chemical parameter (e.g. Br/Cl and I/Cl; Table 2), suggesting NaCl and CaCl<sub>2</sub>-rich fluid inclusions had comparable ranges of U concentration.

5.4. Origin of fluid inclusion excess <sup>40</sup>Ar.

The maximum <sup>40</sup>Ar/<sup>36</sup>Ar values obtained for each sample (considered most representative of the brine inclusions), vary from 307 (RBL2Carb) to 460 (H3042-1) (Table 2). The range of <sup>40</sup>Ar/<sup>36</sup>Ar is typical of upper crustal sedimentary formation waters and K-poor sedimentary rocks (Kendrick et al., 2002a,b; Ozima and Podosek, 2002; Kendrick and Burnard, 2013). The preserved sedimentary succession in the Athabasca Basin has a

maximum depositional age of 1.76 Ga and is characterised by a very low abundance of K (0.1-0.4 wt.%). In contrast, the Archean to Paleoproterozoic basement underlying the Athabasca Basin was >200 Ma old at the ~1.6-1.5 Ga time of mineralisation, and its much higher K content of 2-7 wt.% (Kister et al., 2006 and references therein) suggests it would have been an important reservoir of radiogenic <sup>40</sup>Ar at the time of mineralisation.

Deep percolation of the NaCl and CaCl<sub>2</sub>-rich brines hundreds of meters below the basal unconformity of the Athabasca Basin, into the basement, is demonstrated by the intense alteration of the basement rocks into an illite-sudoite-dravite assemblage (Mercadier et al., 2010). Therefore, the basement is suggested as the most probable source of the excess  ${}^{40}Ar$ component in the ore fluids.

However, basinal fluids could potentially have retained the relatively low, near atmospheric  ${}^{40}$ Ar/ ${}^{36}$ Ar ratios of <450 (Fig. 6), during fluid interaction with the basement, if the amount of <sup>40</sup>Ar derived from the basement was small compared to the amount of atmospheric noble gas derived from eroded sedimentary material. Limited acquisition of <sup>40</sup>Ar is favoured by the low temperature of the fluids which is indicated to have been ~150 to ~200°C by fluid inclusion studies (Derome et al., 2005). The dominantly atmospheric origin of noble gases in solution is consistent with the subaerial origin of the brines, indicated by the halogen data, and it is in stark contrast to ore fluids with <sup>40</sup>Ar/<sup>36</sup>Ar of >30,000 that were interpreted to have been derived from deep within Proterozoic basement lithologies in Australian ore deposits (Kendrick et al., 2011a,b).

### 5.5. Origin of fluid inclusion non-radiogenic isotopes (<sup>36</sup>Ar, <sup>84</sup>Kr, <sup>129</sup>Xe) composition

The lowest calculated fluid inclusion  ${}^{36}$ Ar concentration of  $0.5 \times 10^{-10}$  mol g<sup>-1</sup> (Table 2) is slightly lower than that of modern air saturated meteoric water ( $0.7 \times 10^{-10}$  mol g<sup>-1</sup> at 0°C; Ozima and Podosek, 2002). However, it is considerably higher than calculated for a brine formed by subaerial evaporation of seawater with 30 wt.% salts (bittern brine), which would have a  ${}^{36}$ Ar concentration of only  $0.07 \times 10^{-10}$  mol g<sup>-1</sup> at 25°C (Smith and Kennedy, 1983). The presence of 'excess  ${}^{36}$ Ar', that is greater than can be explained by air saturation, could be partly explained if modern air contaminants trapped in empty fluid inclusions have mixed with noble gases in the fluid inclusions of our samples. However, we do not favour this as a major process because  $F({}^{129}Xe/{}^{36}Ar)_{air}$  is not strongly correlated with  ${}^{40}Ar/{}^{36}Ar$ , and all the fluid inclusion analyses have a range of  $F({}^{129}Xe/{}^{36}Ar)_{air}$  and  $F({}^{84}Kr/{}^{36}Ar)_{air}$  that are distinct from the composition of air (Fig. 6).

Higher than air-saturation <sup>36</sup>Ar concentrations and  $F(^{129}Xe/^{36}Ar)_{air}$  and  $F(^{84}Kr/^{36}Ar)_{air}$ intermediate of air and air-saturated water (Fig. 6) seem to be a real feature of many ore fluids (Kendrick et al., 2011c; Kendrick and Burnard, 2013). Considering ore fluids are chemically reactive solutions (as attested by alteration minerals, see Section 2.3), and atmospheric noble gases can have moderately high abundances in sedimentary rocks (see Table 3 in Podosek et al., 1980; Matsuda and Nagao, 1986; Matsubara et al., 1991; Matsuda and Matsubara, 1989; Ozima and Podosek, 2002), we suggest the fluid inclusions noble gas abundance pattern and high <sup>36</sup>Ar concentration could be explained by fluid interaction with sedimentary minerals. Atmospheric noble gases are adsorbed onto sedimentary materials during deposition, with a portion of the gases trapped in the mineral and a portion released to fluids during diagenesis (e.g. Podosek et al., 1980). Furthermore, atmospheric noble gases adsorbed onto sedimentary materials may be released to the fluid more easily than excess <sup>40</sup>Ar present inside mineral lattices. Therefore, we believe the measured  $F({}^{84}Kr/{}^{36}Ar)_{air}$  and  $F({}^{129}Xe/{}^{36}Ar)_{air}$  values could be fairly representative of NaCl and CaCl<sub>2</sub> fluid inclusions (and minor secondary fluid inclusions), and that these values can be explained by preferential acquisition of <sup>36</sup>Ar><sup>84</sup>Kr><sup>129</sup>Xe from the sedimentary rocks. The heavier noble gases (Xe and Kr) are even more enriched in sediments, than Ar, but are preferentially retained by the rock because they have lower diffusion coefficients and higher adsorption coefficients (Ozima and Podosek, 2002).

#### 5.6. Implications for uranium mineralisation

The mechanisms for reduction of U(VI) (the soluble form) to U(IV) and  $UO_2$ deposition is one of the least understood aspects in the genesis of unconformity-related U deposits. Sangely et al. (2007) demonstrated much of the bitumen in these deposits has  $\delta^{13}$ C of -31 to -49 ‰ that favours an inorganic origin and, together with petrographic observations, casts further doubt on the role of these organic materials. The local occurrence of coprecipitated UO<sub>2</sub> and bitumen has been used to suggest reduction of U(VI) by interaction between U-bearing brines and organic matter including bitumen with  $\delta^{13}C$  of -26.8  $\pm$  0.6 ‰ which may represent kerogen produced by cyanobacteria (Alexandre et al., 2006). It is not possible to exclude the possibility that organic matter could have been locally important for U reduction. However, the low I/Cl ratios of the U-related ore fluids from the four deposits included in this study indicate the brines investigated in this study have not interacted with biogenic organic matter.

The involvement of a basement-derived fluid has been suggested on the basis of stable (O, H) isotope composition of the alteration minerals such as illite, sudoite and dravite  $(\delta^{18}O_{minerals} = 7 \text{ to } 13 \text{ }\% \text{ and } \delta D_{minerals} = -50 \text{ to } -100 \text{ }\%; \text{ Wilson and Kyser } 1987; \text{ Kotzer and}$ Kyser 1995; Alexandre et al., 2005; Cloutier et al., 2009). These authors suggested mixing between a basement-derived (reducing) fluid and basin-derived brines could be responsible for part of the observed alteration and could have triggered UO2 deposition. However, this view was recently challenged by O isotopic analysis of quartz and dolomite cements which points to isotopic equilibrium between the cement-forming fluids (i.e. NaCl- and CaCl<sub>2</sub>-rich brines trapped as fluid inclusions) and the alteration minerals, meaning quartz and dolomite cements as well as illite, sudoite and dravite may have formed from similar NaCl- and CaCl<sub>2</sub>rich brines (Richard et al., 2013b). The noble gas and halogen data presented in this study, suggest the fluid inclusions have an overwhelming origin (e.g. >90 %) from evolved basinal brines produced by subaerial evaporation of Proterozoic seawater. However, it is not possible to completely preclude the input of a volumetrically minor basement fluid (or gas), because if such a fluid had a low noble gas concentration and negligible salinity, its geochemical signature would be overwhelmed by the more voluminous saline fluids documented here.

Our favoured model for mineralisation is that interaction between brines and graphitic metapelites in the basement generated volumetrically minor mobile reductants, such as  $CH_4$ ,  $H_2$  or  $H_2S$  whose interaction with the brines could have triggered UO<sub>2</sub> deposition (Hoeve and Sibbald, 1978; Bray et al., 1988). The possibility also exists that deeply-sourced gases migrated along the faults controlling the U ore, these faults being deeply rooted in the graphitic metapelites (Györfi et al., 2007). Although direct evidence for such processes are currently lacking, such reactions are consistent with the presence of trace  $CH_4$  and  $H_2$  within fluid inclusions (Derome et al., 2003b) and the generally homogenous basinal signature of halogen and noble gas composition of the U-bearing brines documented in this study. Finally, the current dataset stresses the potential implication of fluid-gas interaction for localising the U deposition and encourages further experimental and analytical work on the origin and reactivity of such gases.

#### 6. Conclusion

Analysis of noble gases (Ar, Kr, Xe) and halogens (Cl, Br, I) in U-rich NaCl- and CaCl<sub>2</sub>-rich fluid inclusions from quartz and dolomite veins and breccia cements associated with alteration and mineralisation in four different Proterozoic unconformity-related U deposits in the Athabasca Basin suggest the following conclusions:

Halogen (Br/Cl and I/Cl) systematics demonstrate unambiguously that U-rich NaCland CaCl<sub>2</sub>-rich brines (~25-35 wt.% salts) preserved in fluid inclusions formed by subaerial evaporation of seawater beyond the point of halite saturation. In contrast, the brines K/Cl and Ca/Cl ratios are largely rock-buffered.

The fluid inclusions have <sup>40</sup>Ar/<sup>36</sup>Ar values that are only slightly higher than the modern atmospheric value. This is consistent with the halogen data and supports a subaerial origin from sedimentary formation waters. The fluid inclusions have higher <sup>36</sup>Ar, <sup>84</sup>Kr and <sup>129</sup>Xe concentrations than can be explained by air saturation, probably reflecting the acquisition of excess noble gases through interaction with clastic sediments that contained trapped atmospheric components. In contrast, the fluids acquired only minor excess <sup>40</sup>Ar during interaction with basement rocks.

Finally, the new data put further constraints on the mechanism of UO<sub>2</sub> deposition. The I/Cl data do not favour interaction between brines and organic matter as an important mechanism for precipitating U mineralisation. Together the halogen and noble gas data suggest the fluid inclusions are overwhelmingly dominated by fluids of a surficial origin and indicate that if a basement-derived fluid localised mineralisation it was volumetrically minor. We favour the hypothesis that a volumetrically minor gas phase with low noble gas and halogen concentration (e.g. CH<sub>4</sub>, H<sub>2</sub> or H<sub>2</sub>S) could have played an important role in localising U mineralisation.

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#### 025 **Table captions**

64 65 Table 1: Location, petrography and fluid inclusion mean salinity ( $\pm 2\sigma$ ) of the studied samples. U/C: unconformity; FI: fluid inclusions; *N*: number fluid inclusions analysed per sample by microthermometry. Only brine inclusions were considered (not scarce low-salinity secondary inclusions). Following the nomenclature in Derome et al. (2005), inclusions showing a halite cube at room temperature are noted Lwh' or Lwh according to the nature of their last phase to melt (ice or hydrohalite respectively) during low temperature microthermometric runs and according to their mode of homogenization (by halite or vapor disappearance respectively). Inclusions with no halite cube whose last phase to melt is hydrohalite are noted Lw2. Inclusions with no halite cube whose last phase to melt is ice are noted Lw2 (-60°C < *Tm ice* < -30°C), or Lw1 (-30°C < *Tm ice* < -15°C), where *Tm ice* is the temperature for ice melting. Mineral abbreviations: Qtz: quartz; Dol: dolomite; Su: sudoite; Hem: haematite; Drv: dravite. "bleached" stands for samples altered to illite-sudoite-dravite±quartz assemblages.

Table 2: Summary of fluid inclusion noble gas, halogen and cation data for quartz and dolomite samples in unconformity-related uranium deposits, (Athabasca Basin), determined by *in vacuo* crushing. All ratios are molar. <sup>40</sup> Ar<sub>E</sub>: age-corrected excess <sup>40</sup>Ar (<sup>40</sup>Ar<sub>E</sub> = <sup>40</sup>Ar<sub>total</sub> - <sup>40</sup>Ar<sub>R</sub> - <sup>40</sup>Ar<sub>A</sub>), <sup>40</sup>Ar<sub>R</sub>: radiogenic <sup>40</sup>Ar produced in situ since 1.6 Ga., <sup>40</sup>Ar<sub>A</sub>: atmospheric <sup>40</sup>Ar = 296 × <sup>36</sup>Ar. Representative Ar concentrations have been calculated from mean salinity (Table 1). Fluid inclusion  $F(^{84}Kr/^{36}Ar)_{air}$  versus  $F(^{129}Xe/^{36}Ar)_{air}$  are defined as  $F(X/^{36}Ar)_{air} = (X/^{36}Ar)_{sample}/(X/^{36}Ar)_{air}$ ; reference values for air from Ozima and Podosek (2002). nd.: not determined. dupl.: duplicate analysis.

64 65 \* These samples have age-corrected (1.6 Ga)  ${}^{40}$ Ar/ ${}^{36}$ Ar values lower than the modern atmosphere value of 296, meaning  ${}^{40}$ Ar<sub>E</sub>/Cl and [ ${}^{40}$ Ar<sub>E</sub>] values cannot be calculated for these data.

Table 3: Summary of fluid inclusion noble gas, halogen and cation data for quartz and dolomite samples in unconformity-related uranium deposits, (Athabasca Basin), determined by fusion. All ratios are molar. <sup>40</sup> Ar<sub>E</sub>: age-corrected excess <sup>40</sup>Ar (<sup>40</sup>Ar<sub>E</sub> =  $^{40}$ Ar<sub>total</sub> -  $^{40}$ Ar<sub>R</sub> -  $^{40}$ Ar<sub>A</sub>),  $^{40}$ Ar<sub>R</sub>: radiogenic <sup>40</sup>Ar produced in situ since 1.6 Ga.,  $^{40}$ Ar<sub>A</sub>: atmospheric <sup>40</sup>Ar = 296 ×  $^{36}$ Ar. Fluid inclusion F(<sup>84</sup>Kr/<sup>36</sup>Ar)<sub>air</sub> versus F(<sup>129</sup>Xe/<sup>36</sup>Ar)<sub>air</sub> are defined as F(X/<sup>36</sup>Ar)<sub>air</sub> = (X/<sup>36</sup>Ar)<sub>sample</sub>/(X/<sup>36</sup>Ar)<sub>air</sub>; reference values for air from Ozima and Podosek (2002). \* Sample ES287-10 has K/Cl of ~6.5 meaning fusion released a lot of K and Ar from mineral impurities (see Section 4) and the agecorrected (1.6 Ga) <sup>40</sup>Ar/<sup>36</sup>Ar cannot be calculated. \*\* Sample H3042-1 has a measured  $^{40}$ Ar/<sup>36</sup>Ar of 291, lower than the modern atmosphere value of 296, meaning an age cannot be calculated.

#### **Figure captions**

Figure 1: (A) Simplified geological map of the Athabasca Basin and its underlying basement, with location of the four studied unconformity-related U deposits (modified from Jefferson et al., 2007 and Card et al., 2007). WMTZ: Wollaston-Mudjatik transition zone. (B) Simplified mineral paragenesis for unconformity-related U deposits in the Athabasca Basin. The types of alteration (E: early diagenetic, hydrothermal and late meteoric), are indicated. Hydrothermal alteration and mineralisation spanned a period of <150 Ma between 1670 and 1500 Ma (see Section 2.3. for details) and late meteoric</li>

events are indicated to be younger than 400 Ma (Mercadier et al., 2011b). The alteration assemblage contemporaneous of the primary U mineralisation is composed of, illite, sudoite (Mg-chlorite), dravite (Mg-tourmaline), euhedral quartz, dolomite, haematite and bitumen. Compac.: compaction ; Diag.: diagenetic. APS: aluminium phosphate-sulphate (APS) minerals. Modified from Kotzer and Kyser (1995), Derome et al., (2005), Kyser and Cuney (2008) and Mercadier et al. (2012).

Figure 2: Examples of quartz and dolomite veins and breccia cements, hosting fluid inclusions
studied in this work. (A) Quartz vein crosscutting "bleached" gneiss (sample P48-5, PPatch deposit). (B) Dravite + quartz cementing vein in graphite-rich gneiss (samples
H3042-1, Eagle point Deposit). (C) Quartz and dolomite cementing vein crosscutting a
haematite-rich gneiss (sample RBL2, Rabbit Lake deposit). (D) Typical primary twophase (liquid + vapor) fluid inclusions in quartz. (E) Typical three-phase (liquidvapor-halite) fluid inclusions in quartz. Note that similar two-phase and three phase
inclusions were found in both quartz and dolomite. Abbreviations: Qtz: quartz; Dol:
dolomite; Hem: haematite; Drv: dravite; Liq: liquid; Vap: vapour; Hal: halite. Further
illustration of fluid inclusions and quartz and dolomite veins and breccia cements from
McArthur River, Rabbit Lake, P-Patch and Eagle Point deposits, can be found in Pagel
et al., 1980; Kotzer and Kyser, 1995; Derome et al., 2005; Richard et al., 2010, 2011,
2012, 2013a).

Figure 3: (A) Fluid inclusion Br/Cl versus I/Cl molar ratios determined by *in vacuo* crushing.
Typical 1σ internal uncertainties on Br/Cl and I/Cl values lie within the size of
symbols. The composition of the seawater evaporation trajectory (SET); halite; fluids
that dissolve evaporites; marine pore fluids; mid-continent MVT and magmatic fluids

<sup>2</sup>1076

are based on Zherebstova and Volkova, (1966); Holser (1979); Böhlke and Irwin 1100 1 <sup>2</sup>1101 3 (1992b); Banks et al. (2000); Kendrick et al. (2001; 2002; 2011c); Fehn et al. (2006); <sup>4</sup><sub>5</sub>1102 Muramatsu et al. (2007); the complete list of references for all fields are given in 6 71103 8 9 10 Kendrick and Burnard (2013). Marine pore fluids are characterised by organic Br and I. Assuming an initial seawater Br/Cl value, most marine pore fluids acquire organic 11 121105 Br and I in a ratio of 0.5–2.5 (Kendrick et al., 2011c). (B) Fluid inclusion Br/Cl versus 13 14 15 1106 Cl concentration determined by in vacuo crushing. Br/Cl values are the average Br/Cl 16 171107 among successive extractions for each sample (error bars indicate minimum and 18 <sup>19</sup>201108 maximum values). Cl concentrations have been deduced from the mean salinity ( $\pm 2\sigma$ ) <sup>21</sup> 22<sup>1</sup>109 of the studied samples (as expressed in wt.% eq. NaCl). SET: seawater evaporation 22110) 23 241110 25 26 271111 trajectory after Fontes and Matray (1993). Minerals precipitating during evaporation of seawater: G: gypsum, H: halite, E: epsomite, S: sylvite, C: carnallite, B: bischofite. 28 291112 (C) Comparison between average Br/Cl values determined in this study by *in vacuo* 30 <sup>31</sup><sub>32</sub>1113 crushing (full symbols) and fusion (crossed full symbol symbols), and Br/Cl values 33 341114 obtained on the same samples by crush-leach and ion chromatography (Richard et al., 35 <sup>36</sup>1115 37 2011).

<sup>41</sup>1117 42 Figure 4: Fluid inclusion K/Cl versus Ca/Cl molar ratios determined by in vacuo crushing. 43 441118 The approximate fields for NaCl-rich and CaCl<sub>2</sub>-rich fluid inclusion end-members are 461119 shown for comparison, based on K, Ca (and other cations) determination from LA-48 49</sub>1120 ICP-MS analyses and Cl determination from microthermometry (Richard et al., 2010; 511121 Derome et al., 2005).

561123 Figure 5: (A) Fluid inclusion K/Cl versus Br/Cl molar ratios determined by in vacuo crushing. <sup>58</sup>1124 (B) Fluid inclusion Ca/Cl versus Br/Cl molar ratios determined by in vacuo crushing.

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Sample	Depth below	Description			FI freq	uencies	3		FI relative ab	oundances (%)	Salinity	$\pm 2\sigma$
	U/C (m)		N	laCl-ri	ich	CaC	l <sub>2</sub> -rich	Ν	NaCl-rich	CaCl <sub>2</sub> -rich	mean	(%)
			Lw1	Lw2	Lwh	Lw'	Lwh'				(wt% eq. NaC	l)
Eagle Pou	nt deposit - Qua	artz										
ES287-10	259	Qtz vein in bleached pelitic gneiss	6	17	0	0	0	23	100	0	28.5	10
H3042-1	263	Qtz+drv vein in graphite-rich gneiss	2	5	0	15	7	29	24	76	30.3	25
P-Patch d	eposit - Quartz											
P48-2	39	Otz vein in chloritised gneiss	1	14	0	7	4	26	58	42	29.9	15
P48-5	52	Qtz vein in bleached gneiss	3	4	0	12	1	20	35	65	27.0	15
McArthur	River deposit -	Ouartz										
MAC54Q	z 50	Qtz+drv+sulfides cement in Su-rich breccia	0	47	7	0	0	54	100	0	27.5	10
Rabbit Lal	ke deposit - Ou	artz										
RBL2Qz	* ~	Otz+dol vein in hem-rich gneiss	1	46	0	0	0	47	100	0	29.2	10
RBL7Qz	*	Qtz+dol cement in hydraulic megabreccia	0	5	0	2	2	9	56	44	28.9	30
Rabbit Lak	ke deposit - Dol	lomite										
RBL2Carb	, *	Otz+dol vein in hem-rich gneiss	14	25	0	0	0	39	100	0	27.6	5
RBL7Carb	*	Otz+dol cement in hydraulic megabreccia	1	12	0	2	0	15	87	13	28.1	10

Sample	Sample	Extraction	Cl	Br/Cl	L ±	I/Cl	±	K/Cl	±	Ca/Cl ±	U/Cl	เบา	<sup>36</sup> Ar	Age		<sup>40</sup> Ar/ <sup>36</sup> A	r ±	<sup>40</sup> Ar/ <sup>36</sup> A	Ar ±	F( <sup>84</sup> Kr/ <sup>36</sup> Ar) <sub>air</sub>	. ±	F( <sup>129</sup> Xe/ <sup>36</sup> Ar) <sub>air</sub>	±	Cl <sup>/36</sup> Ar ±	<sup>40</sup> Ar <sub>E</sub> /Cl	±	[ <sup>36</sup> Ar]	[ <sup>40</sup> Ar <sub>E</sub> ]
name	weight	Linucuon	mol g <sup>-1</sup>	(×10 <sup>-3</sup>	3)	(×10 <sup>-6</sup>	)					ppm	mol g <sup>-1</sup>	Ma	±	Measured	d d	(1.6 Ga	.)	- ( , , an		- ( )an		$(\times 10^{6})$	(×10 <sup>-6</sup> )		mol g <sup>-1</sup>	mol g <sup>-1</sup>
	(mg)		$(\times 10^{-6})$	(,,10	/	(,,10	,						$(\times 10^{-14})$											((10))	(1.6 Ga)		$(\times 10^{-10})$	$(\times 10^{-8})$
	(8/		(,(10))										(//10 )												(110 01)		(×10 )	(1.6 Ga)
Eagle Point depo	osit - Quar	tz																										
ES287-10	88.7	cr #1	1.3	7.6	0.2	2.0	0.1	0.033	0.002	0.08 0.02	nd	l	5.5	3436	719	350	4	337	4	1.04	0.04	1.19	0.05	23.2 0.3	1.8	0.2	2.1	0.9
	88.7	cr #2	0.6	7.3	0.2	1.9	0.1	0.038	0.001	nd	7.8E-06	9	1.0	3581	865	467	25	428	23	1.20	0.08	1.64	0.19	58.2 3.1	2.3	0.3	0.8	1.1
ES287-10 (dupl.)	49.4	cr #1	1.3	7.3	0.2	2.4	0.1	0.056	0.004	nd	nd	l	5.7	3137	712	369	7	347	7	1.05	0.05	1.35	0.09	22.4 0.5	2.3	0.3	2.2	1.1
H3042-1	95.4	cr #1	1.4	10.4	0.1	2.3	0.1	0.026	0.002	nd	2.5E-06	3	4.7	4022	988	376	4	362	4	1.07	0.04	1.20	0.05	30.3 0.3	2.2	0.1	1.7	1.2
	95.4	cr #2	0.7	10.2	0.2	2.5	0.2	0.040	0.006	0.11 0.03	nd	l	10	3696	1379	318	2	313	2	1.67	0.06	1.89	0.09	6.7 0.1	2.7	0.3	7.8	1.4
	95.4	cr #3	0.6	9.5	0.2	1.8	0.1	0.024	0.002	0.07 0.04	7.1E-06	9	0.8	4160	1035	478	16	449	15	1.30	0.06	2.70	0.15	69.1 2.3	2.2	0.1	0.7	1.2
H3042-1 (dupl.)	36.3	cr #1	1.5	9.3	0.2	2.2	0.1	0.027	0.005	0.18 0.05	nd	l	5.0	3891	1837	371	13	357	13	0.98	0.05	1.06	0.08	29.2 1.0	2.1	0.4	1.8	1.1
P-Patch deposit	- Quartz																											
P48-2	98.2	cr #1	1.6	9.5	0.2	2.6	0.1	0.032	0.001	0.12 0.01	1.6E-05	20	8.2	2460	275	318	1	307	1	1.02	0.04	1.39	0.06	19.7 0.1	0.6	0.1	2.6	0.3
	98.2	cr #2	1.5	9.5	0.2	2.7	0.1	0.031	0.001	0.12 0.01	4.8E-05	58	5.7	2660	384	329	2	315	2	1.02	0.04	1.10	0.05	26.2 0.2	0.7	0.1	1.9	0.4
	98.2	cr #3	1.7	9.5	0.2	2.6	0.1	0.031	0.001	0.10 0.02	2.7E-05	33	4.7	2689	556	343	6	324	5	1.06	0.04	1.08	0.06	36.0 0.7	0.8	0.1	1.4	0.4
	98.2	cr #4	2.1	9.0	0.2	2.5	0.1	0.032	0.001	0.11 0.01	5.4E-05	66	5.8	2395	299	335	2	315	2	1.02	0.04	1.12	0.06	36.5 0.4	0.5	0.1	1.4	0.3
<b>D</b> (0, <b>r</b> )	98.2	cr #5	2.3	9.1	0.2	2.5	0.1	0.029	0.001	0.08 0.01	1.4E-04	169	4.1	2773	289	370	3	341	3	1.00	0.05	1.34	0.13	56.8 0.7	0.8	0.05	0.9	0.4
P48-5	107.5	cr #1	6.0	8.7	0.1	2.9	0.1	0.028	0.001	0.09 0.01	4.1E-06	5	23	3075	226	337	1	324	1	1.07	0.04	1.15	0.05	26.9 0.1	1.1	0.04	1.7	0.5
	107.5	cr #2	2.6	8.5	0.2	2.8	0.1	0.027	0.001	0.08 0.01	4.3E-06	5	5.9	3004	245	360	2	339	2	1.11	0.04	1.37	0.06	44.9 0.4	1.0	0.03	1.0	0.5
	107.5	cr #3	2.2	8.3	0.2	2.8	0.1	0.027	0.001	0.08 0.01	7.7E-06	8	3.8	2866	336	3/1	3	343	3	1.09	0.04	1.63	0.09	57.1 0.6	0.8	0.04	0.8	0.4
	107.5	cr #4	2.4	8.1	0.2	2.2	0.0	0.029	0.001	0.08 0.01	9.4E-06	10	2.8	3154	221	440	4	396	3	1.16	0.04	1.43	0.11	85.7 0.9	1.2	0.03	0.5	0.6
McArthur River a	deposit - Q	Quartz																										
MAC54Qz	103.2	cr #1	1.2	6.6	0.1	8.2	0.2	0.041	0.001	0.13 0.03	2.6E-05	29	8.0	2229	373	314	2	303	2	1.01	0.04	1.33	0.06	15.0 0.1	0.5	0.1	3.1	0.2
	103.2	cr #2	0.9	6.3	0.1	6.7	0.2	0.041	0.001	0.09 0.02	2.9E-05	33	4.7	2193	324	318	2	304	2	1.02	0.04	1.34	0.07	18.4 0.2	0.5	0.1	2.5	0.2
	103.2	cr #3	1.4	6.4	0.1	6.0	0.2	0.045	0.002	0.10 0.01	5.5E-05	61	4.9	1778	432	323	5	300	4	1.02	0.04	1.31	0.07	29.3 0.5	0.1	0.1	1.6	0.1
	103.2	cr #4	0.6	6.3	0.1	4.6	0.2	0.044	0.002	0.13 0.03	8.6E-05	96	1.6	1655	337	327	5	297	4	1.04	0.04	1.33	0.11	38.6 0.6	0.04	0.1	1.2	0.02
Rabbit Lake depo	osit - Qua	rtz																										
RBL2Qz	95.5	cr #1	0.5	6.0	0.1	5.6	0.2	0.048	0.006	0.10 0.06	nd	1	2.7	2185	463	319	1	305	2	1.04	0.04	1.14	0.10	16.7 0.1	0.6	0.04	3.0	0.3
	95.5	cr #2	0.6	5.9	0.1	5.1	0.1	0.056	0.002	0.11 0.03	nd	l	4.8	2344	765	320	5	307	5	1.03	0.04	1.14	0.05	13.2 0.2	0.9	0.4	3.8	0.4
<b>DDI 50</b>	95.5	cr #3	0.9	5.8	0.1	4.8	0.1	0.054	0.002	0.10 0.01	nd	1	3.0	2073	377	338	5	310	5	1.09	0.04	1.46	0.07	29.6 0.5	0.5	0.1	1.7	0.2
RBL7Qz	91.6	cr #1	0.6	8.3	0.1	3.1	0.1	0.033	0.005	0.09 0.05	1.6E-05	19	1.7	1778	584	320	4	299	5	0.97	0.04	1.02	0.04	35.7 0.5	0.1	0.1	1.4	0.1
	91.6	cr #2	0.5	8.4	0.2	3.0	0.1	0.034	0.005	0.06 0.04	3.5E-05	41	0.9	1719	946	332	17	299	16	0.97	0.07	1.17	0.16	54.2 2.7	0.1	0.3	0.9	0.04
	91.6	cr #3	0.3	8.1	0.2	2.9	0.1	0.037	0.002	0.19 0.06	1.0E-04	118	0.4	630	225	309	4	263	5	0.86	0.05	1.88	0.15	69.4 0.8	*		0.7	*
Rabbit Lake depo	osit - Dolo	omite								-		_																
RBL2Carb	132	cr #1	0.4	7.8	0.1	4.8	0.4	0.047	0.004	nd	6.4E-05	72	16	3260	1267	304	2	301	2	1.00	0.07	1.15	0.04	2.7 0.1	2.2	0.5	17.3	1.0
	132	cr #2	0.6	8.0	0.3	3.8	0.3	0.046	0.007	nd	3.7E-05	42	11	2561	2995	304	12	300	12	1.01	0.05	1.19	0.06	5.0 0.2	1.0	2.5	9.5	0.5
	132	cr #3	0.3	7.9	0.2	4.0	0.1	0.035	0.004	nd	3.5E-05	39	4.2	2989	936	310	2	305	2	1.02	0.04	1.31	0.07	7.8 0.1	1.2	0.2	6.0	0.6
RBL7Carb	101.4	cr #1	2.5	8.8	0.2	3.3	0.1	0.027	0.001	nd	1.8E-05	21	25	2487	509	306	1	301	1	1.01	0.04	1.18	0.05	10.2 0.1	0.5	0.1	4.7	0.3
	101.4	cr #2	1.0	8.7	0.2	2.8	0.1	0.029	0.001	nd	7.5E-05	86	4.7	1178	184	302	1	292	1	0.96	0.07	1.12	0.06	20.9 0.2	*		2.3	*
	101.4	cr #3	0.5	8.4	0.2	3.0	0.1	0.017	0.005	nd	1.3E-04	148	2.0	2994	1627	319	4	311	5	0.97	0.04	1.23	0.19	26.8 0.4	0.6	0.1	1.8	0.3

Table 2

Sample	Sample	C1	Br/Cl	±	I/C1	±	K/Cl	±	U/C1	<sup>36</sup> Ar	Age	$^{40}$ Ar/ $^{36}$ Aı	±	<sup>40</sup> Ar/ <sup>36</sup> Ar	± F	( <sup>84</sup> Kr/ <sup>36</sup> Ar) <sub>air</sub>	±	F( <sup>129</sup> Xe/ <sup>36</sup> Ar) <sub>air</sub>	±
name	weight	mol g <sup>-1</sup>	(×10 <sup>-3</sup> )		(×10 <sup>-6</sup> )					mol g <sup>-1</sup>	Ma ±	Measured	l	(1.6 Ga)					
	(mg)	(×10 <sup>-6</sup> )								(×10 <sup>-14</sup> )									
Eagle Poin	t deposit ·	- Quartz																	
ES287-10	38.6	0.8	5.7	0.1	7.2	1.5	6.46	0.11	2.2E-03	6.3	1376 28	1543	4	*		1.57	0.12	51.56	2.16
H3042-1	31.8	0.7	6.6	0.1	4.4	0.8	0.28	0.03	1.8E-03	7.6	**	291	5	246	12	1.46	0.12	20.93	1.13
P-Patch de	posit - Qu	uartz																	
P48-2	93.9	3.0	6.5	0.1	3.9	0.4	0.45	0.01	1.2E-02	38.8	961 84	326	1	264	2	1.51	0.20	2.77	0.08
McArthur I	River depo	osit - Quar	tz,																
MAC54Qz	95.7	2.4	5.0	0.04	5.1	0.3	0.28	0.01	1.3E-02	26.0	817 124	314	1	268	2	1.97	0.26	2.59	0.07
Rabbit Lak	e deposit	- Quartz																	
RBL2Qz	72.5	0.8	4.4	0.04	4.5	0.2	0.76	2.01	4.7E-04	6.7	993 175	374	4	221	8	1.86	0.11	14.26	0.61
	33.1	2.2	4.9	0.04	4.1	0.2	0.23	10.32	1.2E-04	14.2	831 438	322	4	258	12	0.92	0.06	2.90	0.19











